



APPLICAZIONI DELLA TEORIA DELLA DECISIONE E DELLA STIMA

Fabrizio LOMBARDINI, Lucio VERRAZZANI

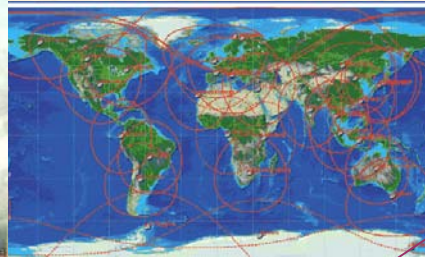
Pisa, 20 Dicembre 2007

<http://www.ing.unipi.it/~d9263/>

Sommario

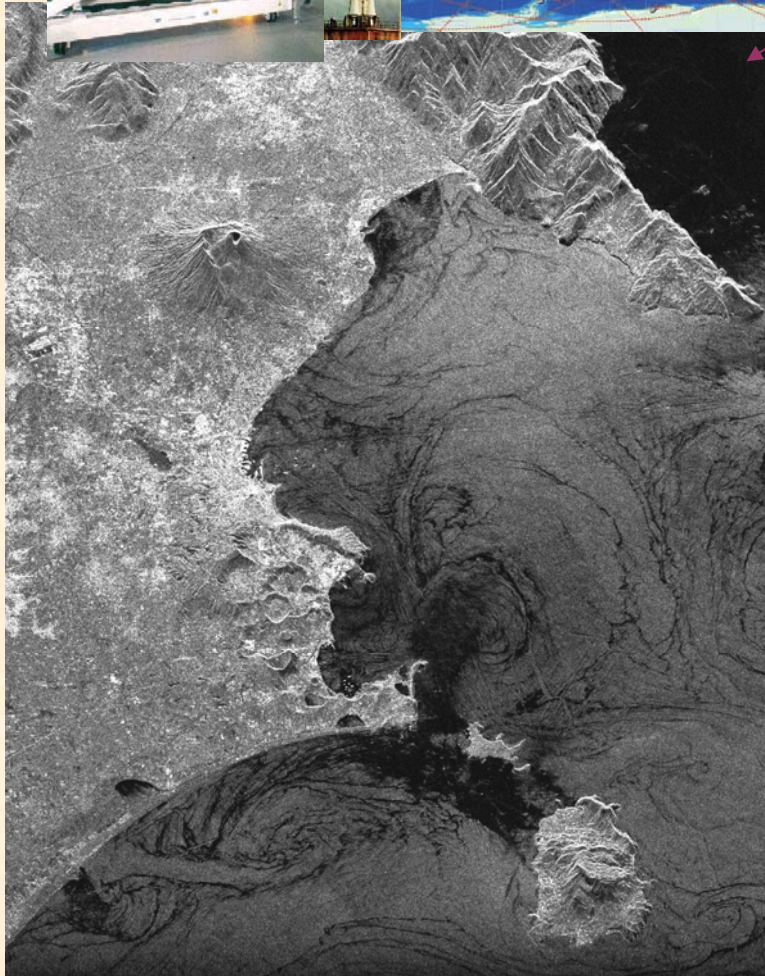
- ④ Array signal processing
- ④ Stima spettrale in rumore moltiplicativo, a super-risoluzione, adattiva
- ④ Stima spettrale spazio-temporale
- ④ Interpolatori e filtri non stazionari
- ④ Bound di Cramér-Rao e di Cramér-Rao ibridi
- ④ Robust signal processing
- ④ Interferometria SAR
- ④ Tomografia radar 3D
- ④ Imaging differenziale
- ④ Studi di sistema per radar satellitari multicanale

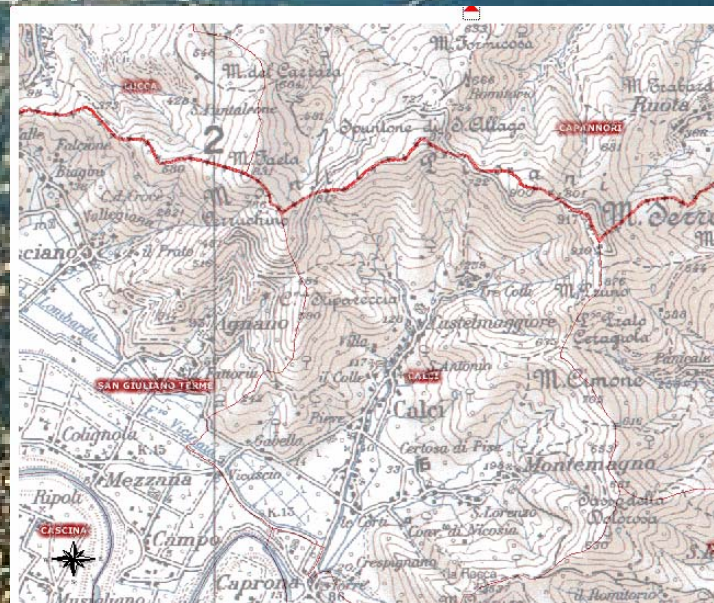
Radar d'immagine (SAR, 2D)



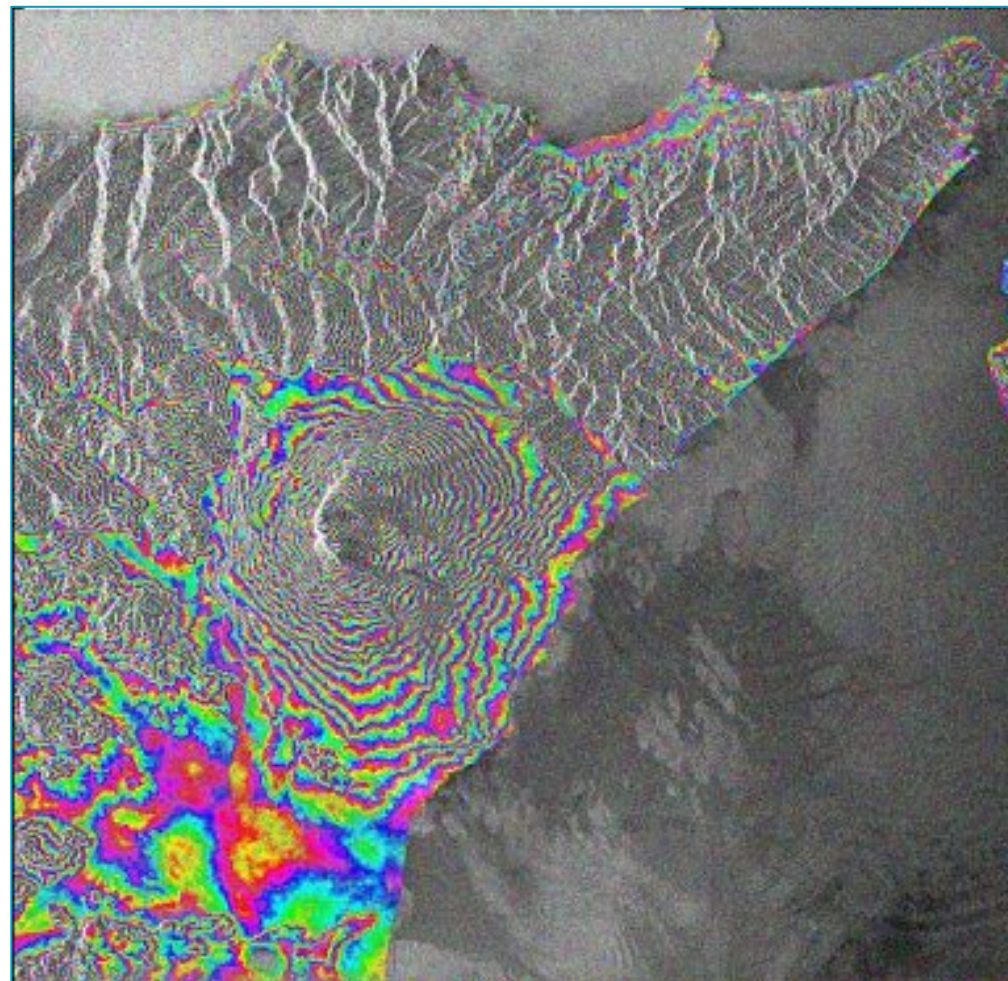
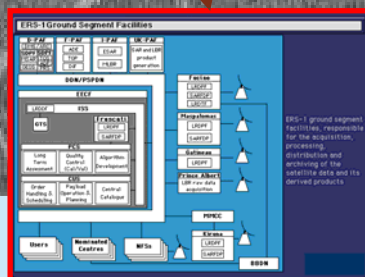
SAR satellitare

SAR aviotrasportato

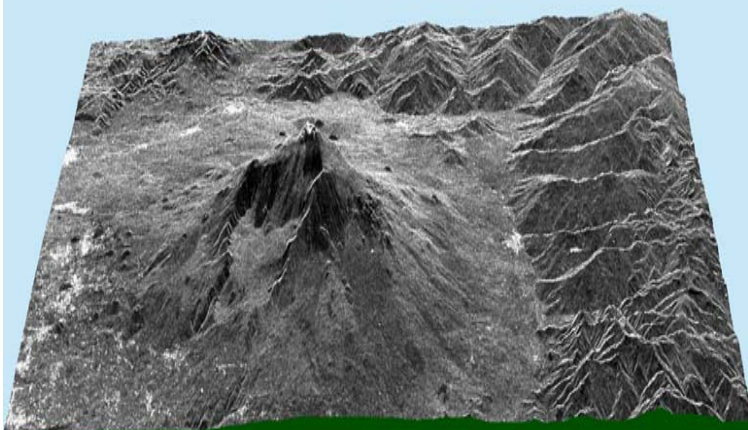




Interferometric SAR – “2.5 D” !

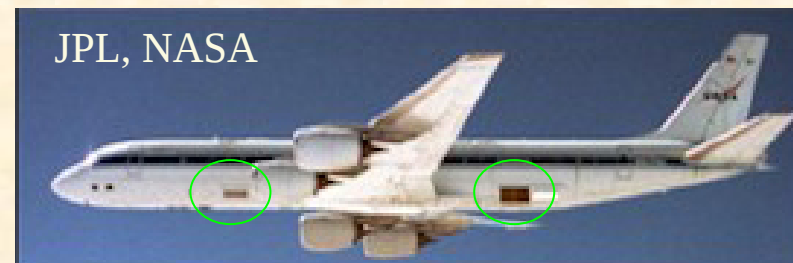
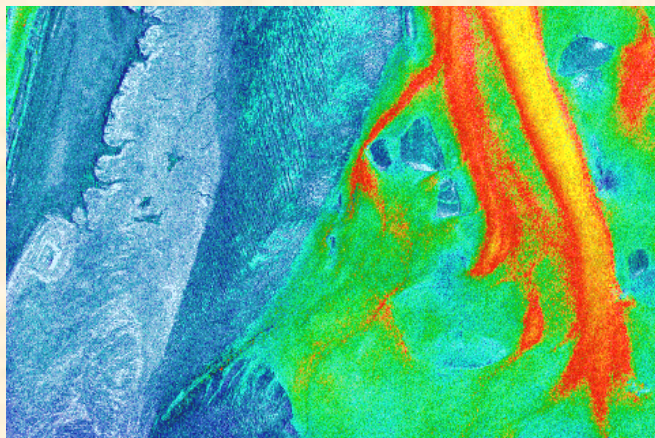


Etna, Sicilia
satellite europeo ERS-1/2

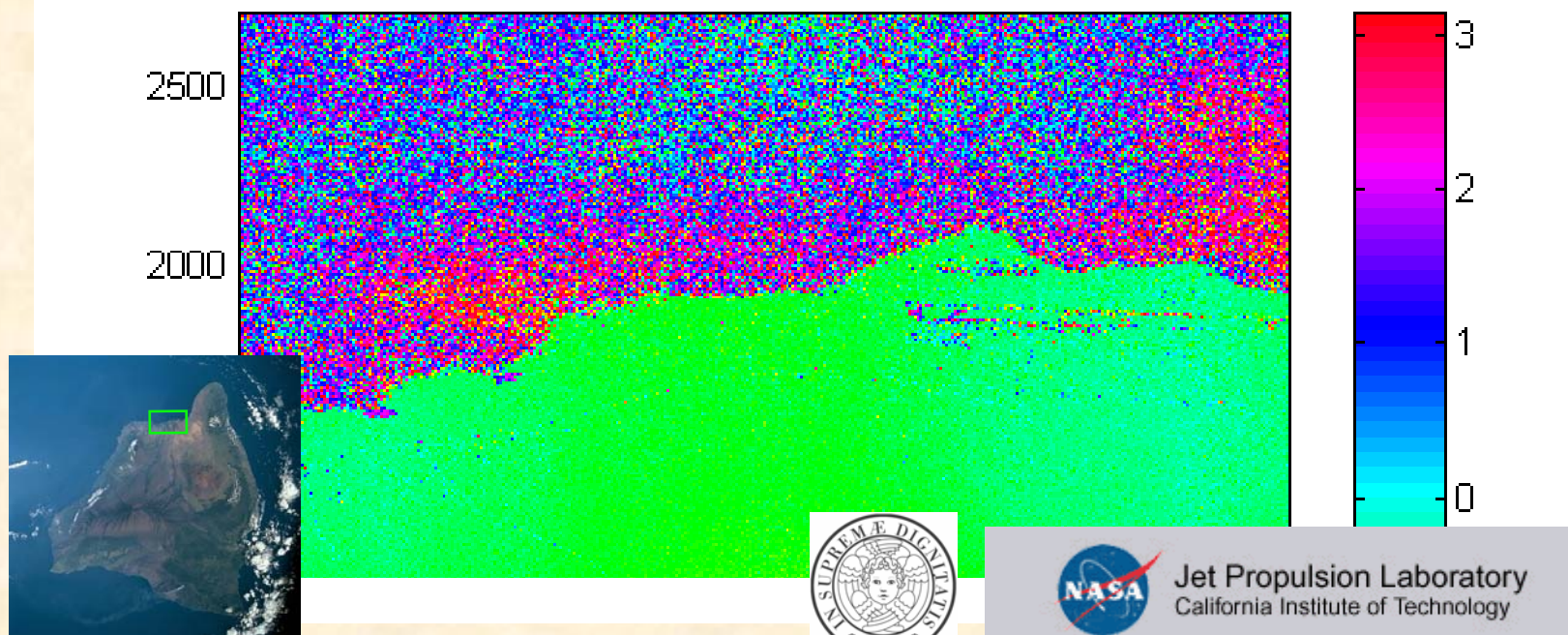




Interferometric SAR – 2D + T!

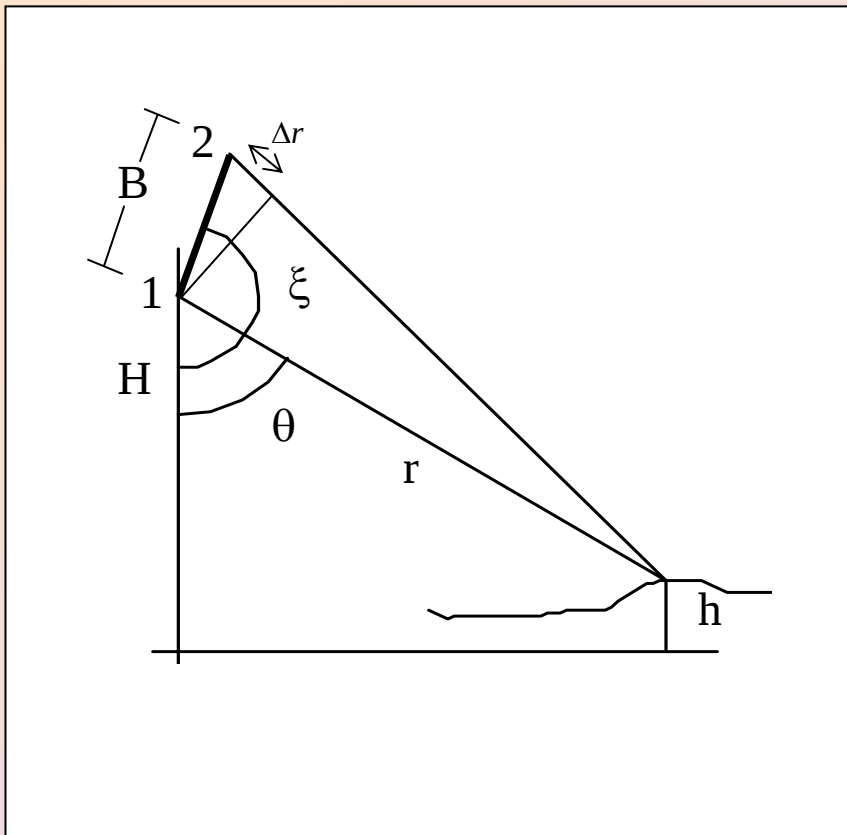


Kohala coast, Big Island, Hawaii

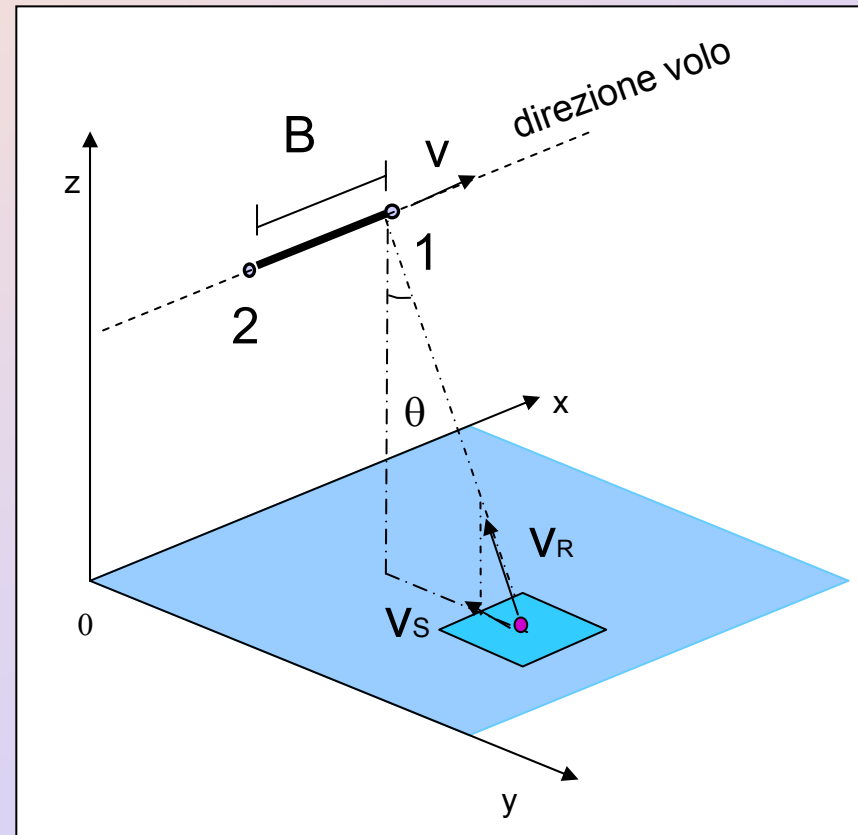


Jet Propulsion Laboratory
California Institute of Technology

Tecnica InSAR e ATI-SAR



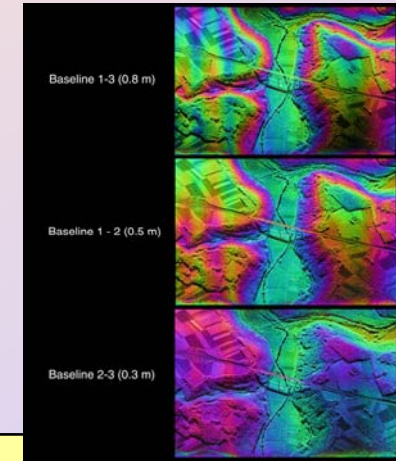
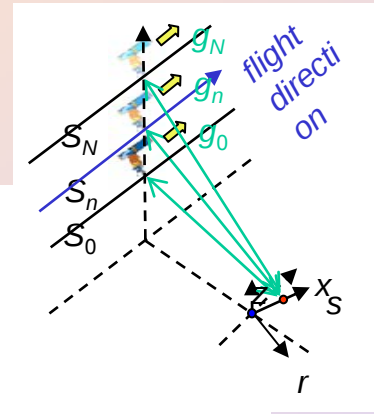
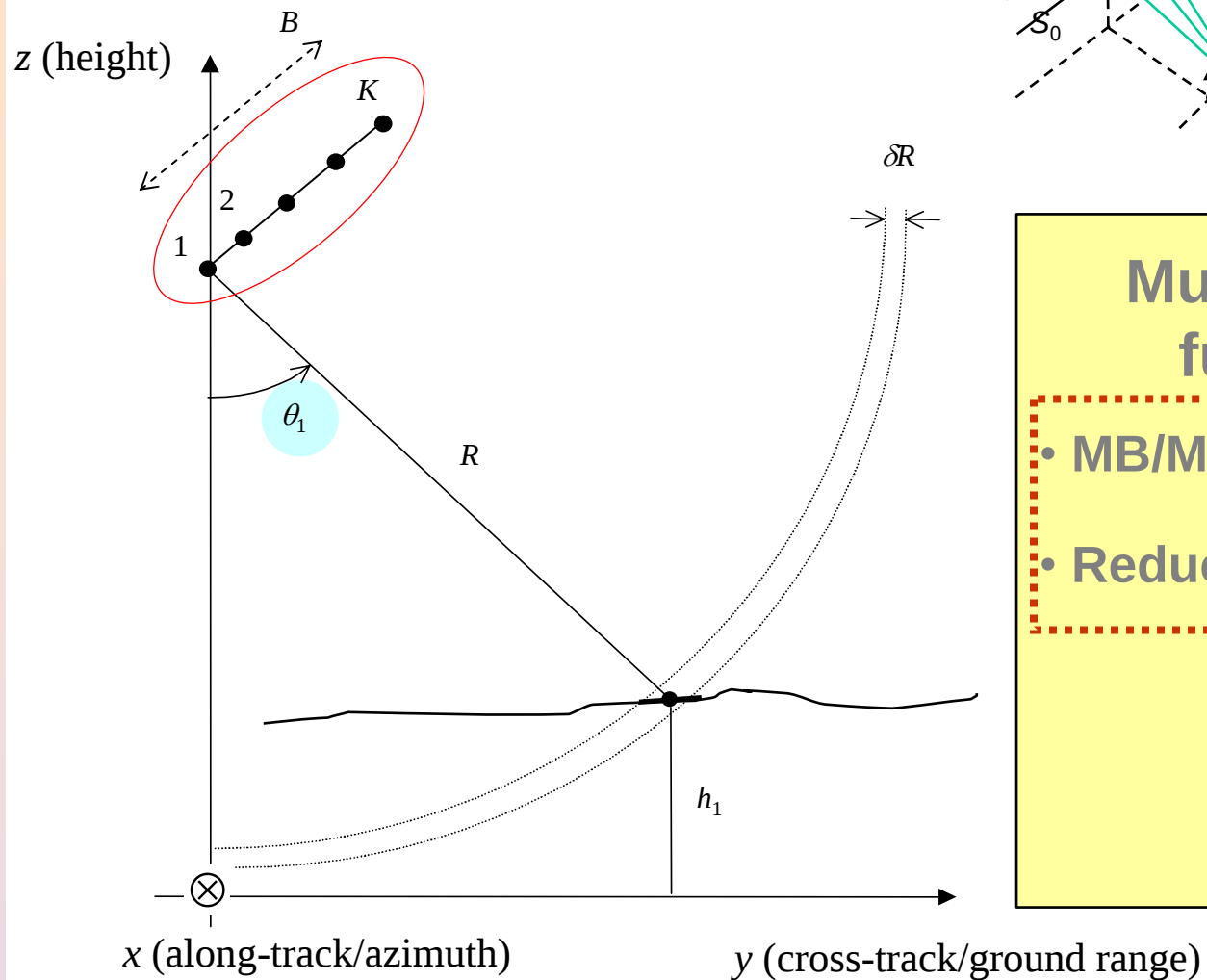
mappe topografiche



mappe correnti marine

Prodotti principali:

Multichannel (MB/MF) InSAR



Multich. improved functionalities:

- MB/MF phase unwrapping
- Reduction of data noise

Array processing

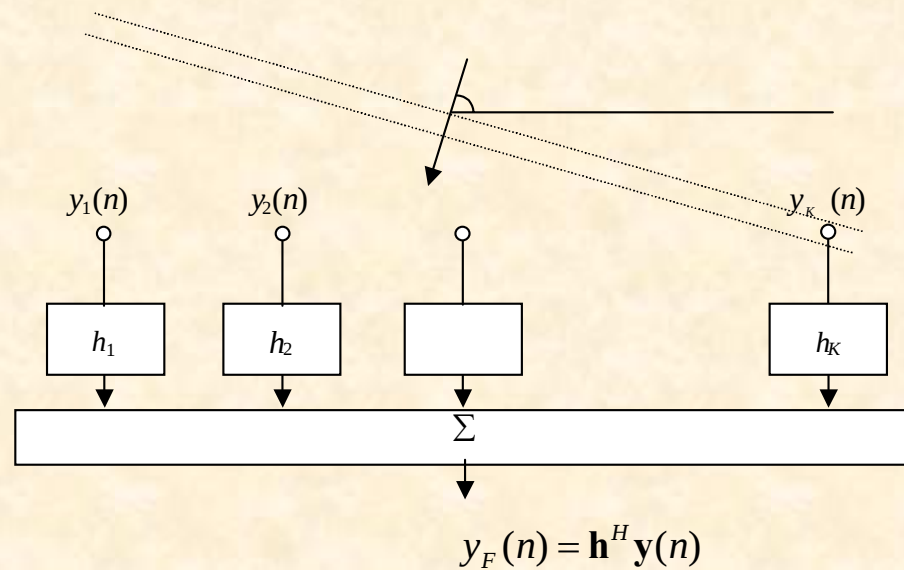
$$\mathbf{y}(n) = [y_1(n) \ y_2(n) \ \dots \ y_K(n)]^T$$

data vector, $n=1,2,\dots,N$ looks

$$\mathbf{a}(\omega) = [1 \ e^{jl_2\omega} \ \dots \ e^{jl_K\omega}]^T$$

steering vector

Spatial frequency!



$$\hat{P}_B(\omega) = \frac{1}{N} \sum_{i=1}^N \left| \underbrace{\frac{1}{K} \mathbf{a}^H(\omega)}_{\mathbf{h}^H} \mathbf{y}(n) \right|^2$$

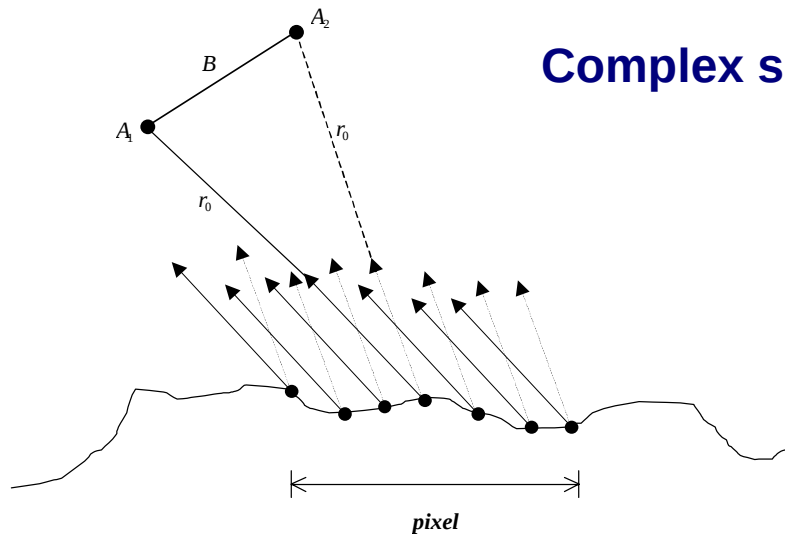
FT

conventional Beamforming
(multilook Periodogram)

$$\hat{P}_B(\omega) = \frac{\mathbf{a}^H(\omega) \hat{\mathbf{R}}_y \mathbf{a}(\omega)}{K^2}$$

Multiplicative noise

11



Complex signal (I & Q) at each pixel of the SAR image

Fully developed speckle: complex circular Gaussian distributed

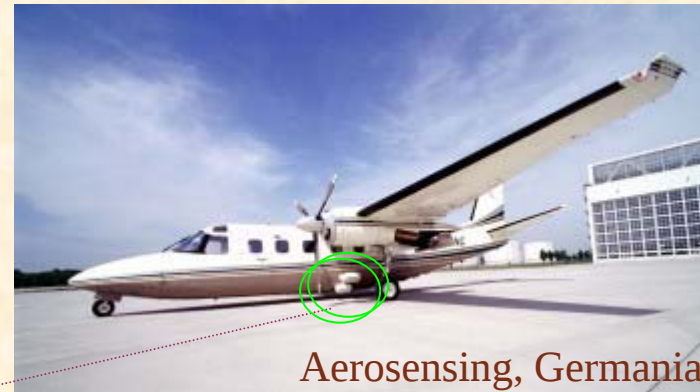
Spatial (baseline) decorrelation: non-line spatial spectrum!

Radar d'immagine (Interferometria, 2.5D)

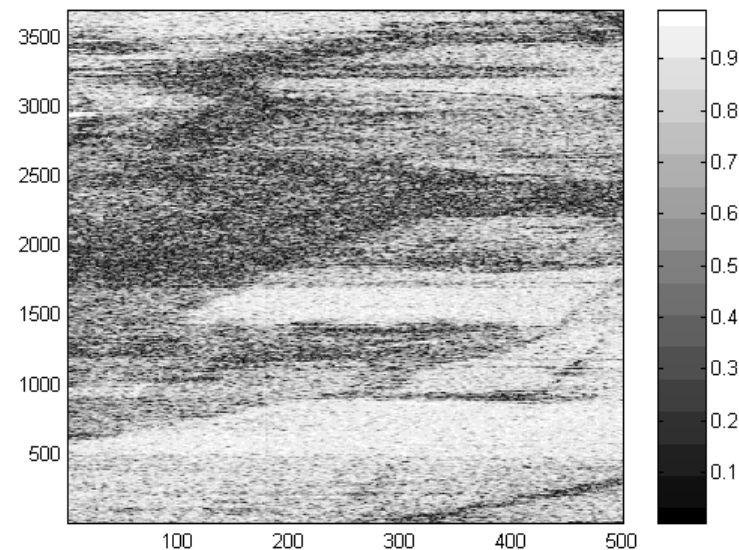
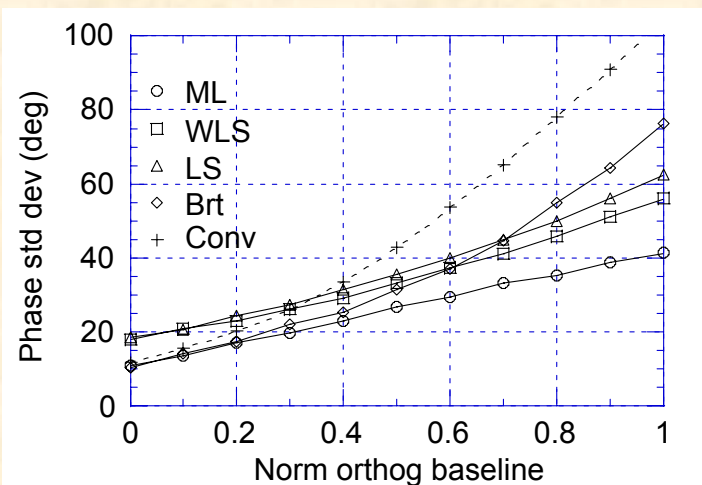
Frequency MLE in multiplicative noise

Stima di fase interferometrica (frequenza spaziale)

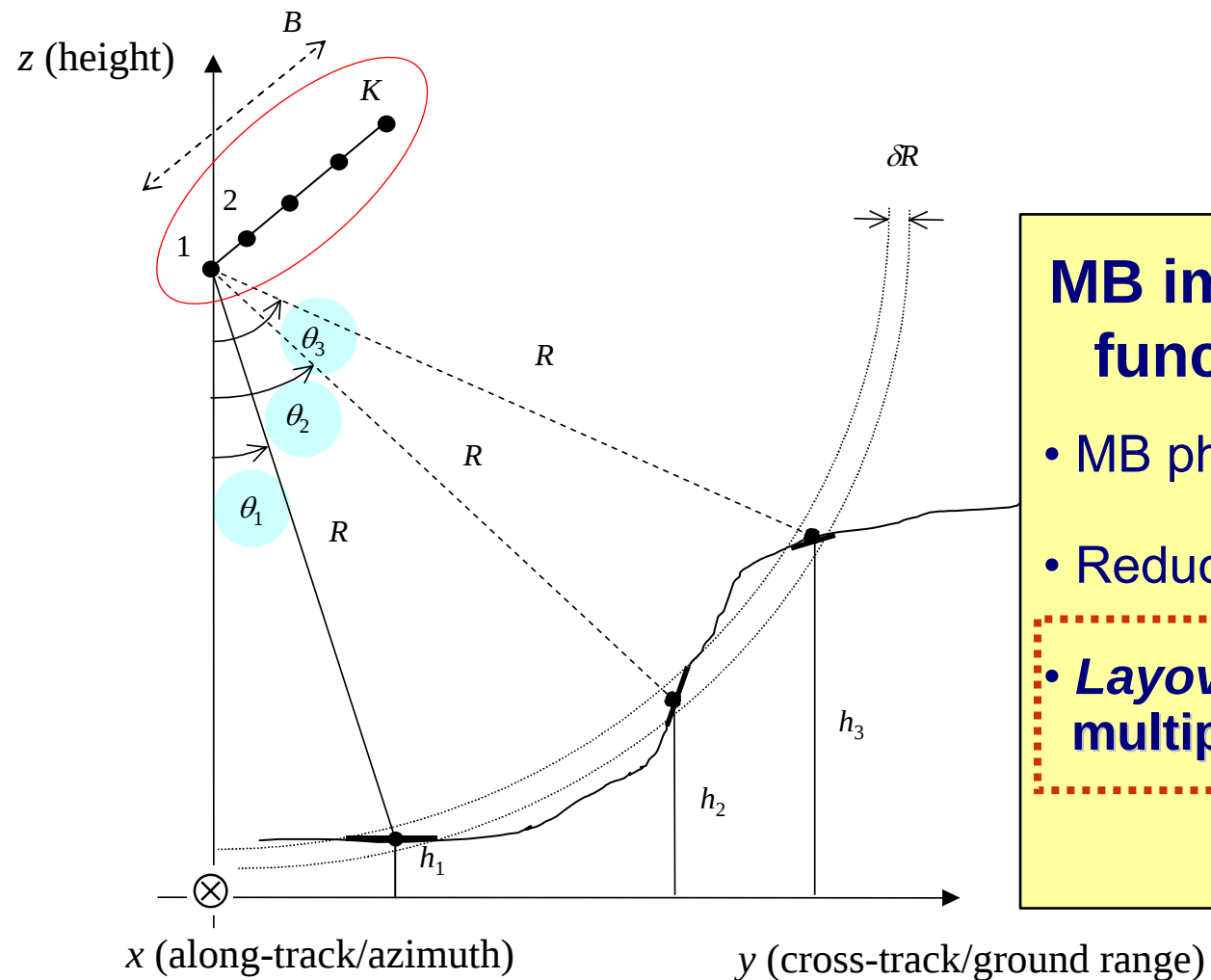
$$J(\varphi) = (\rho_{12} - \rho_{23}\rho_{13}) \operatorname{Re}\{\exp[-j\varphi(1-p)] \sum_{i=1}^N \tilde{P}_1^{*(i)} \tilde{P}_2^{(i)}\} + \\ (\rho_{13} - \rho_{12}\rho_{23}) \operatorname{Re}\{\exp[-j\varphi] \sum_{i=1}^N \tilde{P}_1^{*(i)} \tilde{P}_3^{(i)}\} + \\ (\rho_{23} - \rho_{12}\rho_{13}) \operatorname{Re}\{\exp[-j\varphi p] \sum_{i=1}^N \tilde{P}_2^{*(i)} \tilde{P}_3^{(i)}\}$$



Aerosensing, Germania



OrbiSat
da Amazônia SA



MB improved/new functionalities:

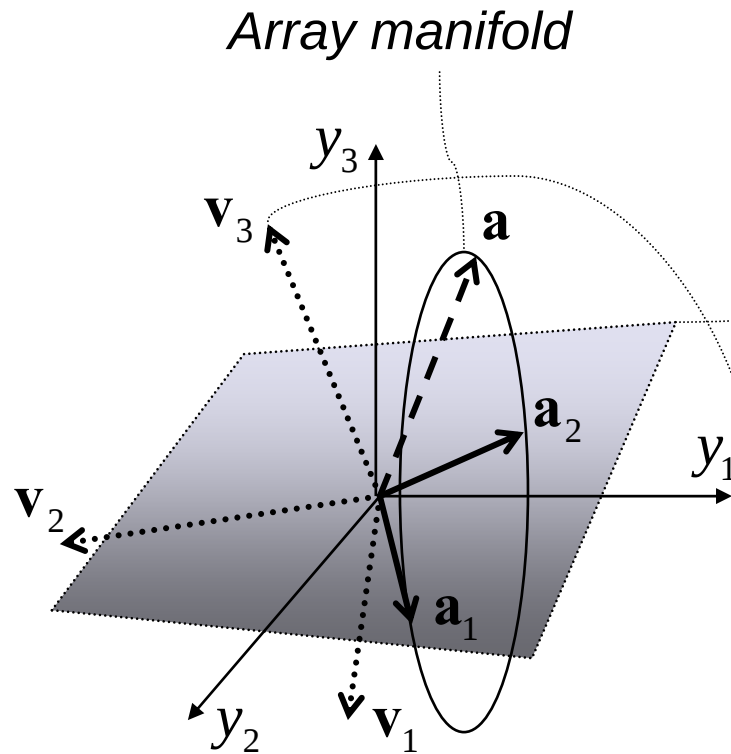
- MB phase unwrapping
- Reduction of data noise
- **Layover overcoming:
multiple DOAs estimation**

Superresolution

14

Zero spatial bandwidth

$$K = 3, N_s = 2 \quad \sigma_v^2 = 0: \quad \mathbf{y}(n) = \alpha_1(n)\mathbf{a}(\varphi_1) + \alpha_2(n)\mathbf{a}(\varphi_2) = \begin{bmatrix} \mathbf{a}_1 & \mathbf{a}_2 \end{bmatrix} \begin{bmatrix} \alpha_1 \\ \alpha_2 \end{bmatrix}$$



Signal subspace

$$N_s, \{\varphi_m\}$$

$$\mathbf{C}_y = \begin{bmatrix} \mathbf{a}_1 & \mathbf{a}_2 \end{bmatrix} \begin{bmatrix} \tau_1 & \tau_2 \end{bmatrix} \begin{bmatrix} -\frac{\mathbf{a}_1^H}{\mathbf{a}_2^H} - \end{bmatrix}$$

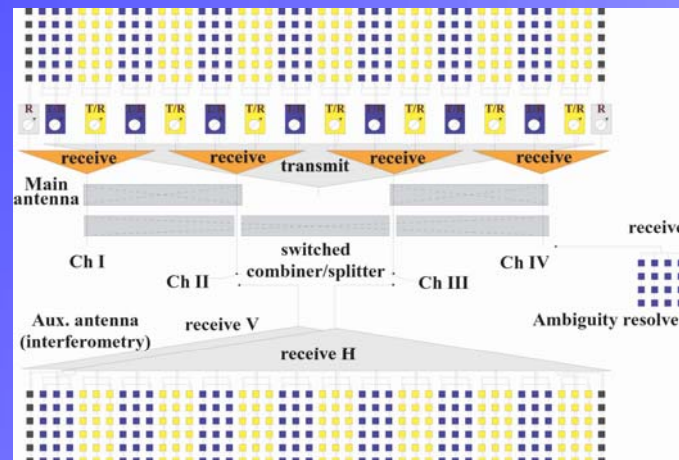
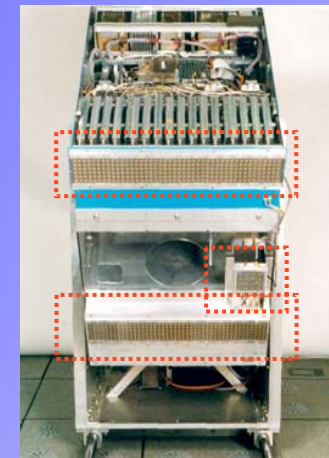
$\text{Range}[\mathbf{C}_y] = \text{Signal subspace}$

“Noise” subspace

Parametric spectral estimation

Airborne experimental results: FGAN-FHR AER-II

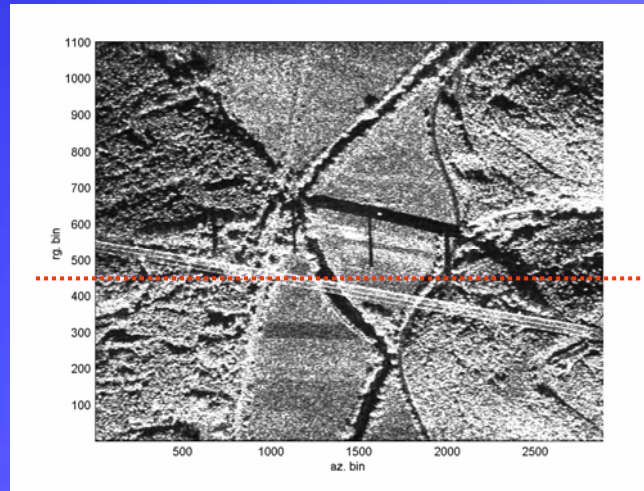
Center frequency	10.0 GHz
Resolution	1 m
Range	20 km
XTI channels	3 parallel receive channels, overall baseline 0.8 m
Antenna	Active phased array, 16 T/R modules + auxiliary antennas
Transmit power	80 W
Polarisation	VV (HH / HV / VH)



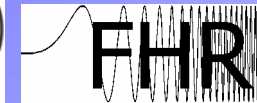
Airborne multiantenna results (FGAN AER-II)



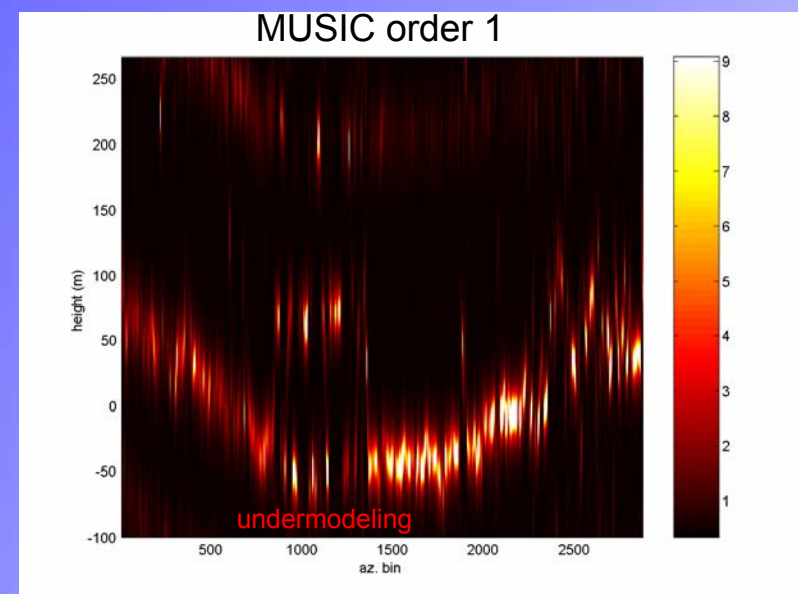
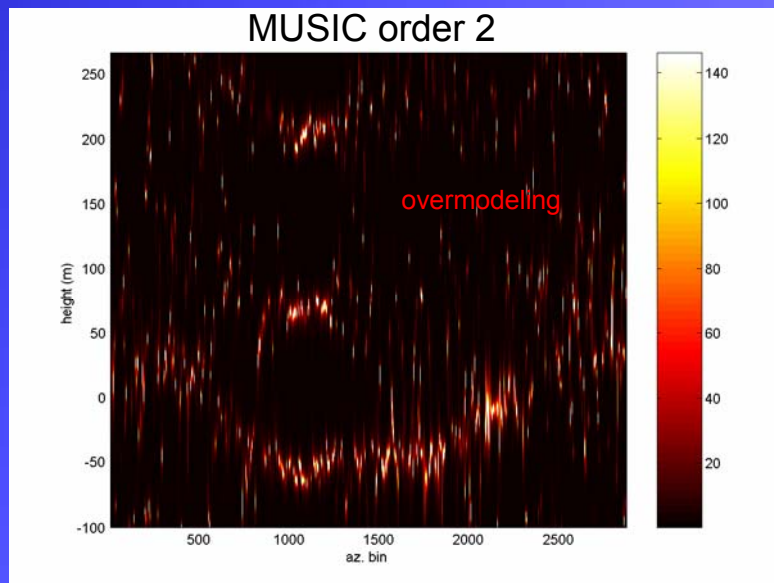
Neckar river valley and highway bridge,
Weitingen, Germany

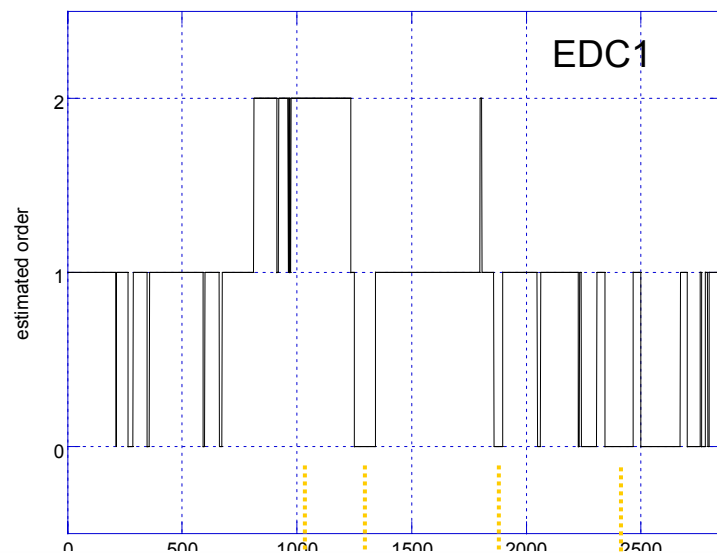


- THERMAL NOISE
EQUALIZATION
- PHASE CALIBRATION
- AMPLITUDE CALIBRATION
- DERAMPING AND ORTHO
BASELINE EVALUATION



*Height
resolution
98 m
Ambiguous
height
370 m*

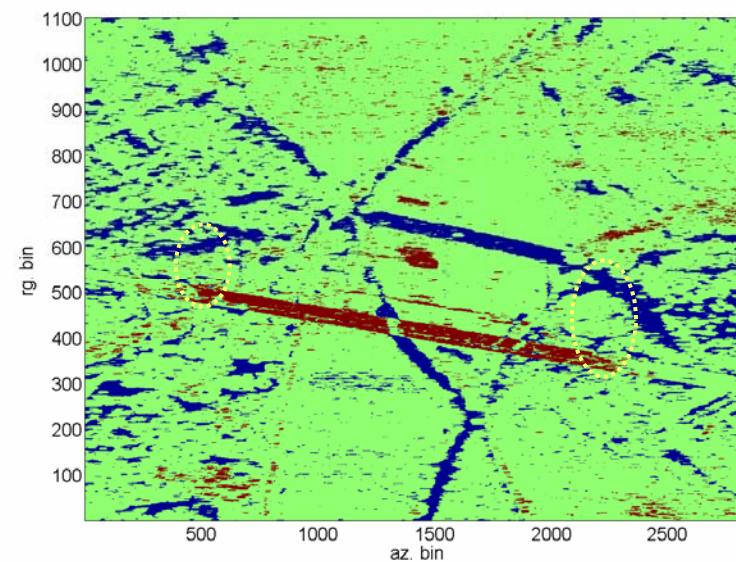
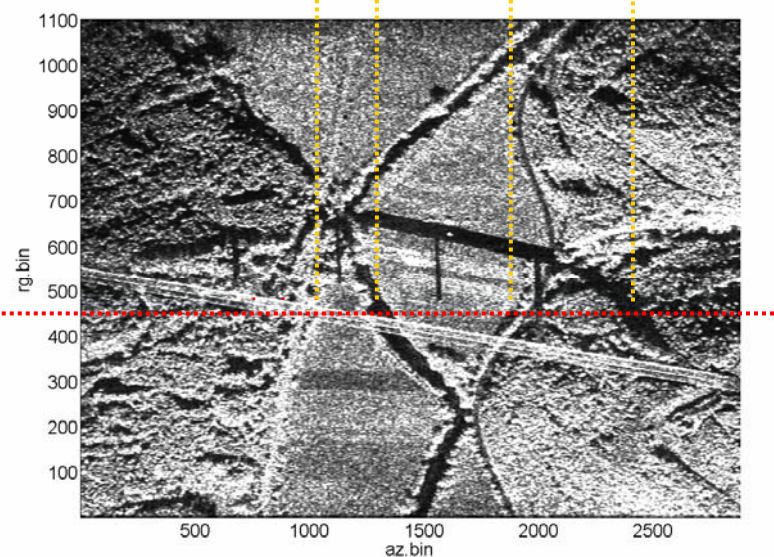




Numero di look
pari a 33 looks,
stabilizzazione
degli autovalori

Prestazione della stima dell'ordine:
90%

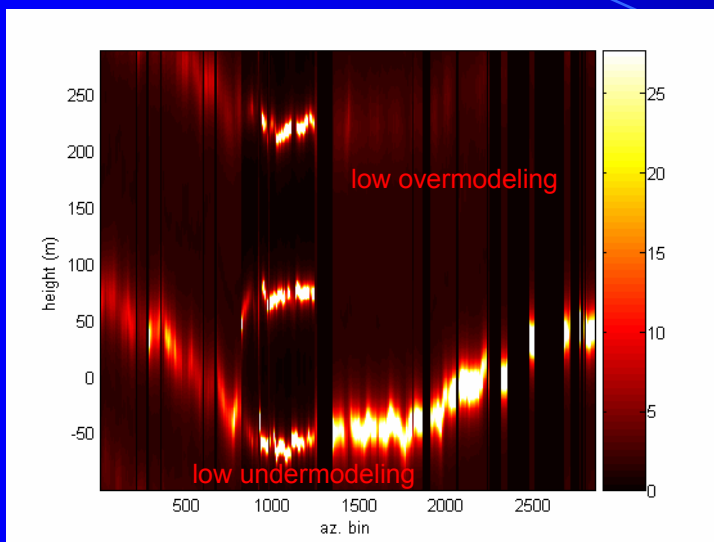
r, g, b: $\hat{N}_s = 2, 1, 0$



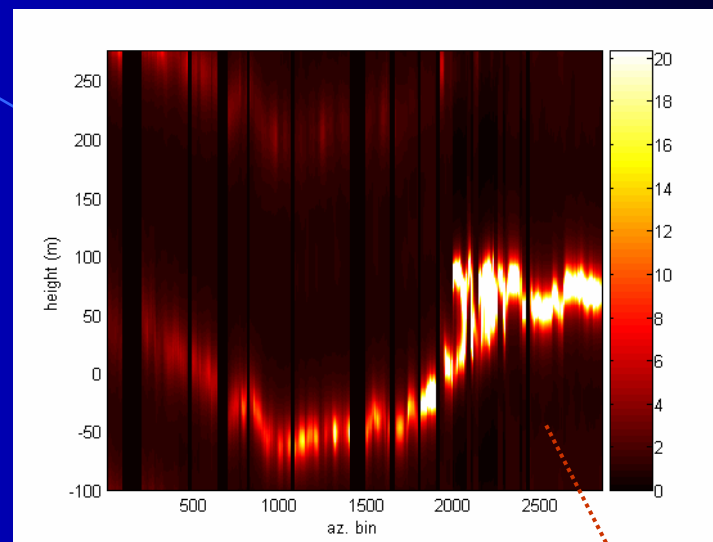


Metodo integrato

EDC-MUSIC

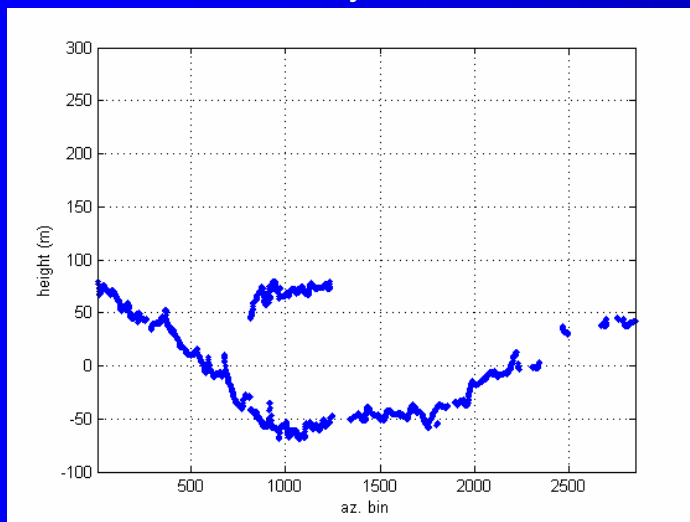


EDC-MUSIC

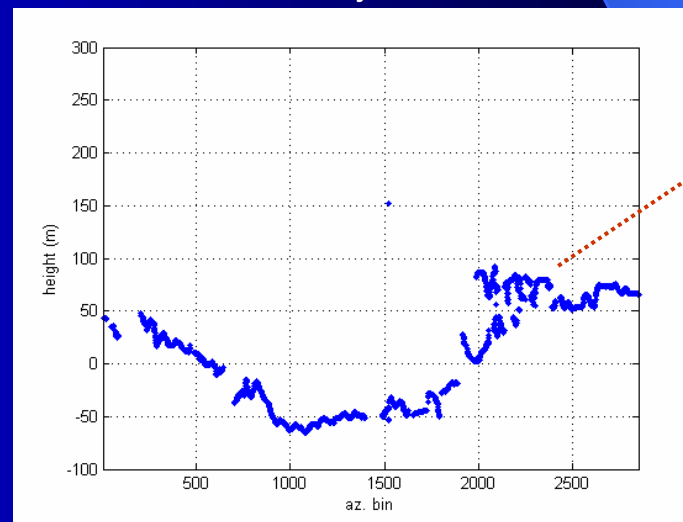


Post-processing: si sfruttano le info a priori sul settore in quota in cui sono contenuti i picchi da MUSIC

Altezze in layover estratte



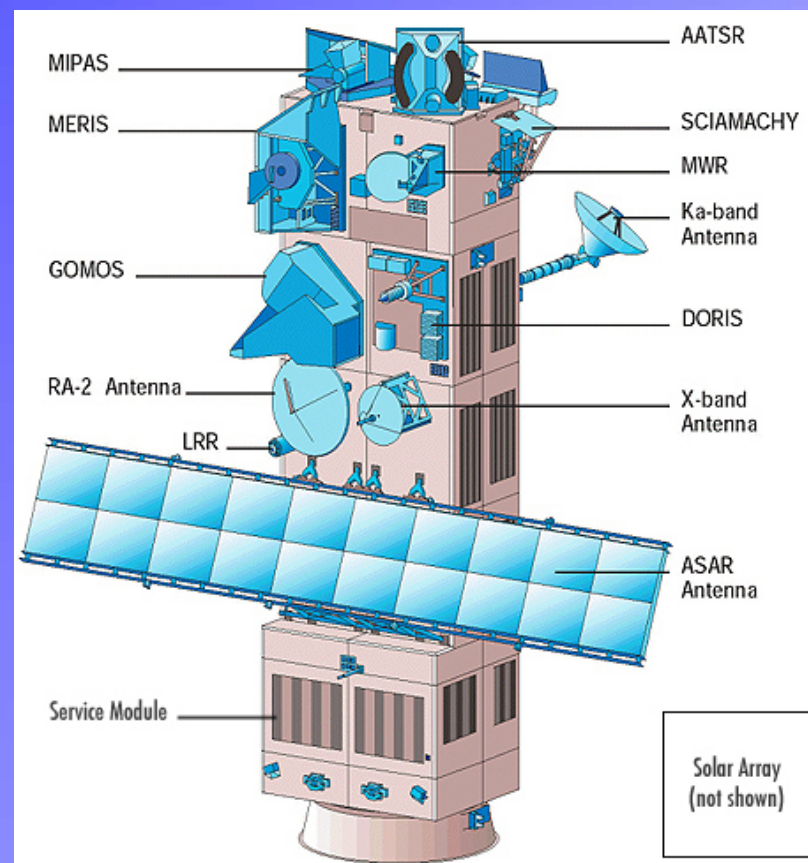
Altezze in layover estratte



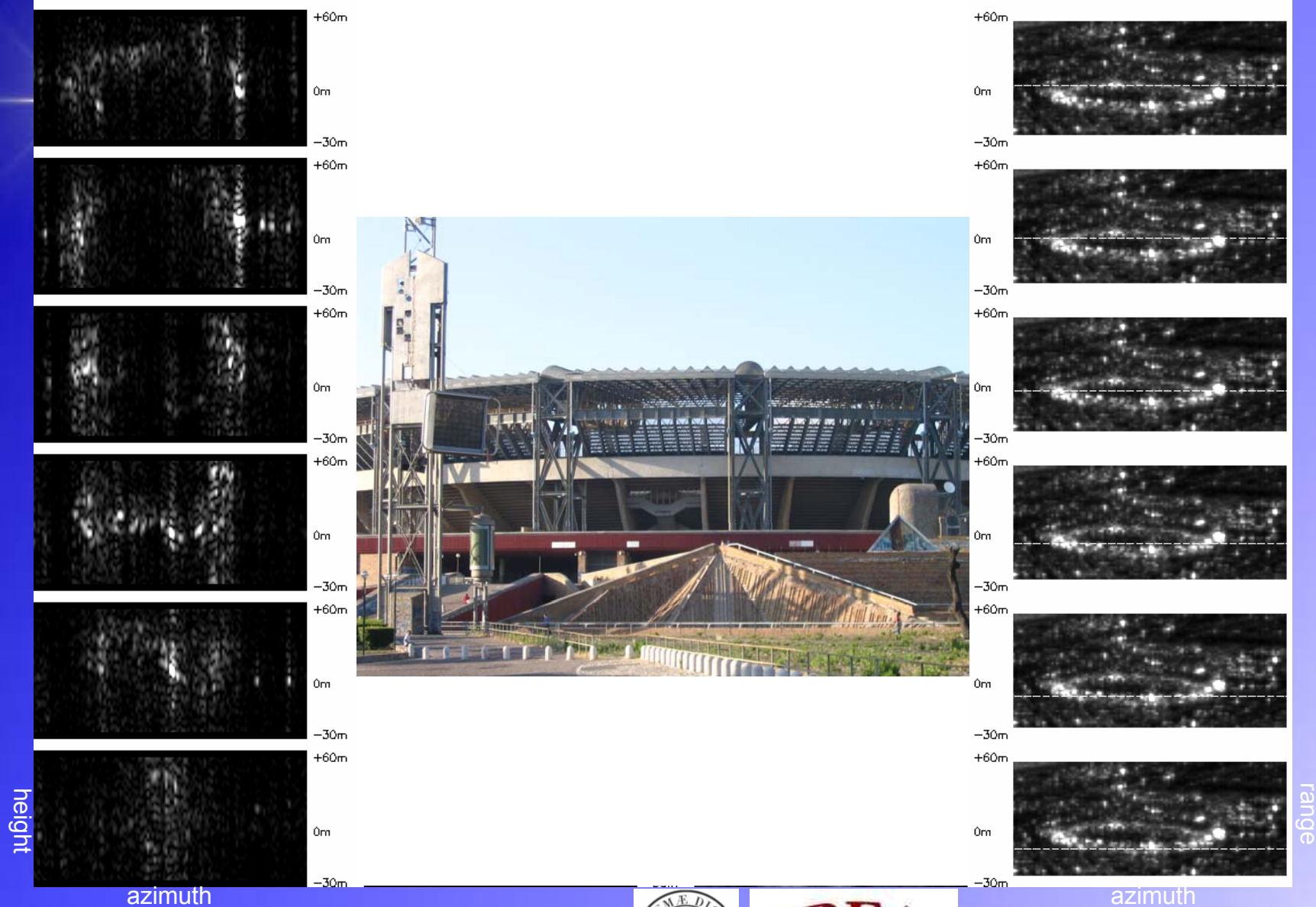
bridge
end

Complex structures/discontinuous surfaces $\rightarrow$ **layover** $\rightarrow$ multicomponent signal





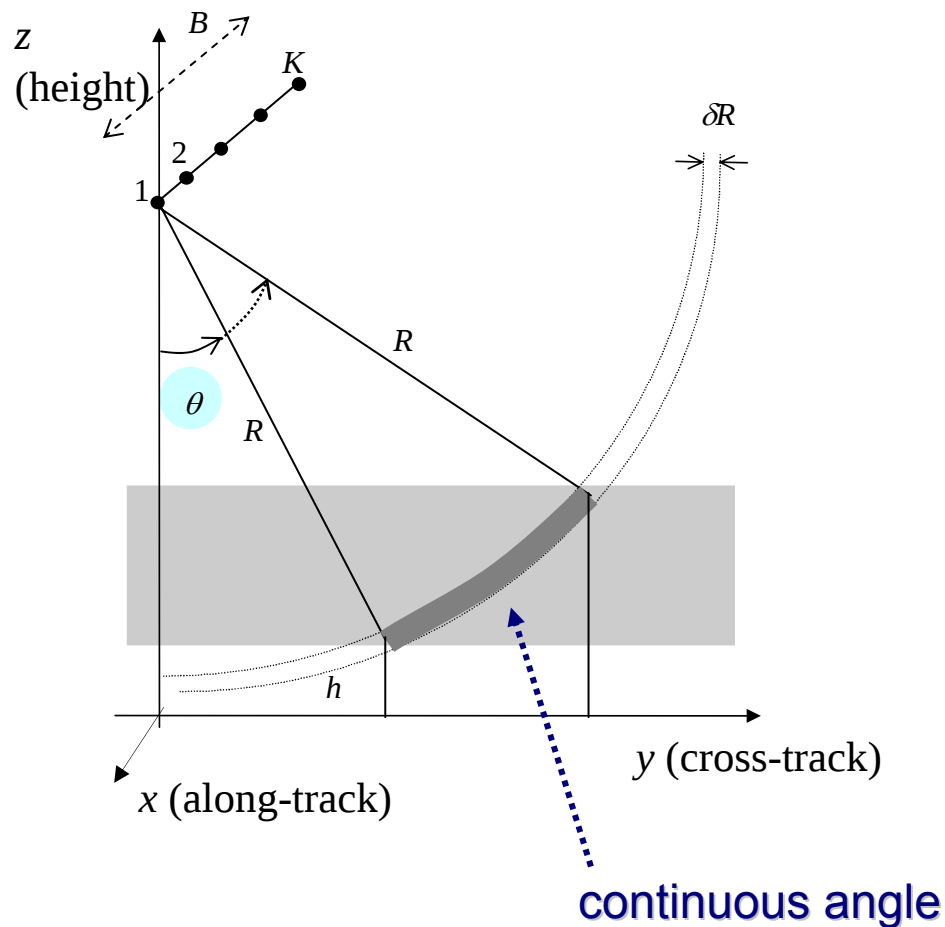
S.Paolo Stadium







Layover: the illuminated 3-D reflectivity $\mathbf{a}(\mathbf{x}, \mathbf{y}, \mathbf{z})$ has been assumed to be concentrated to a 2-D surface



Aim of XTI-SAR:
extract τ_i and θ_i from data

from 2.5-D
to 3-D

**SAR-Tomography
(full 3-D):**
when $\mathbf{a}(\theta)$ is
continuous w.r.t. θ ,
not discrete

“TAC” radar!

Radar d'immagine (Tomografici, Full 3D)



E-SAR **DLR**
L-band, VV

DGPS+INS

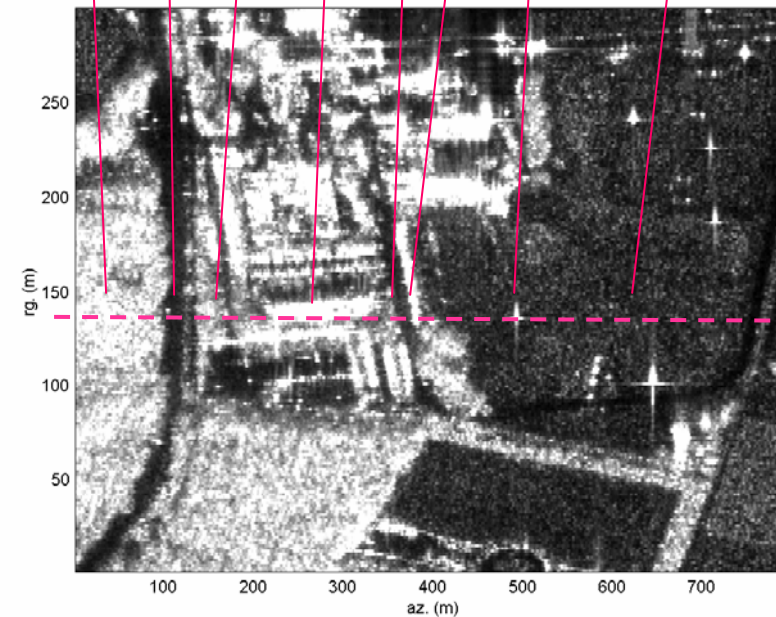
K=14 tracks
overall baseline: L=240 m

site: **Oberpfaffenhofen Germany**

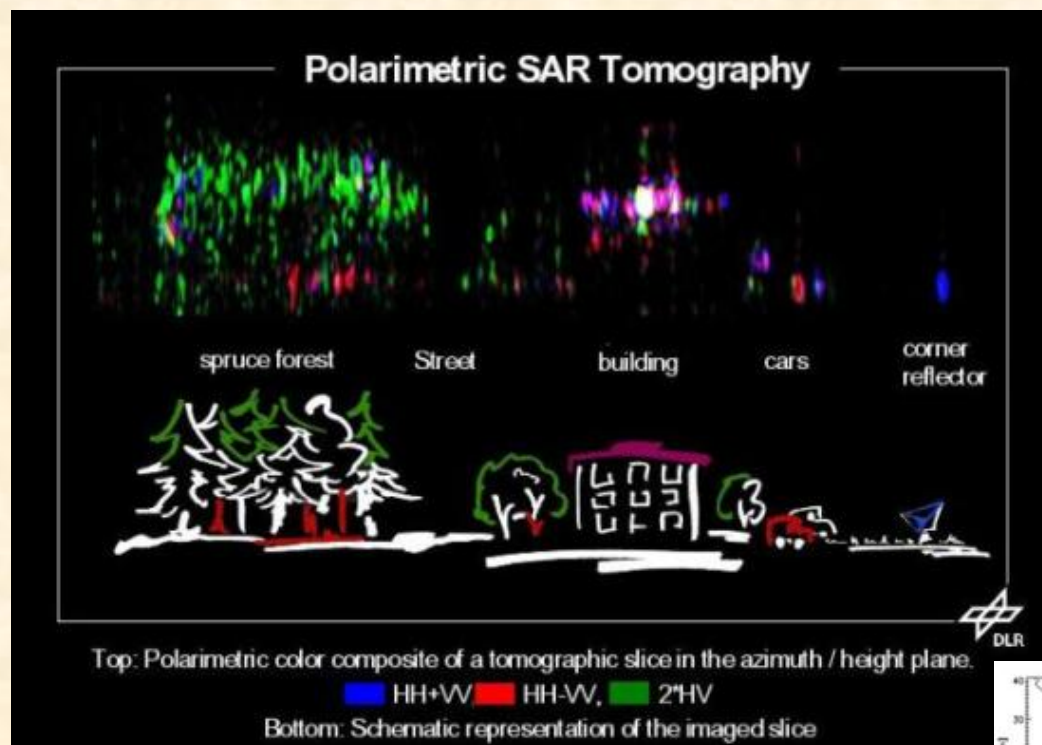
Fourier height resolution: 3 m
unambiguous height range: 35 m

MB phase calibration:
corner reflector

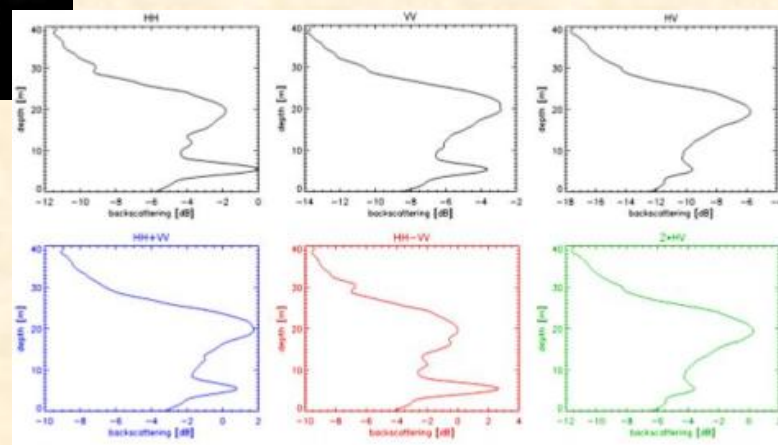
spruce forest bushes bushes corner reflector
street building cars grass land



Radar d'immagine (Tomografici, Full 3D)



Stima spettrale spaziale
non-parametrica



Current Problems of Radar Tomography

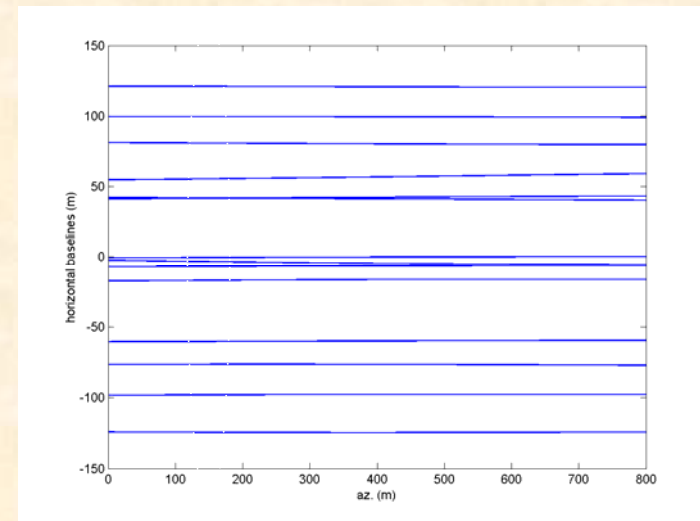
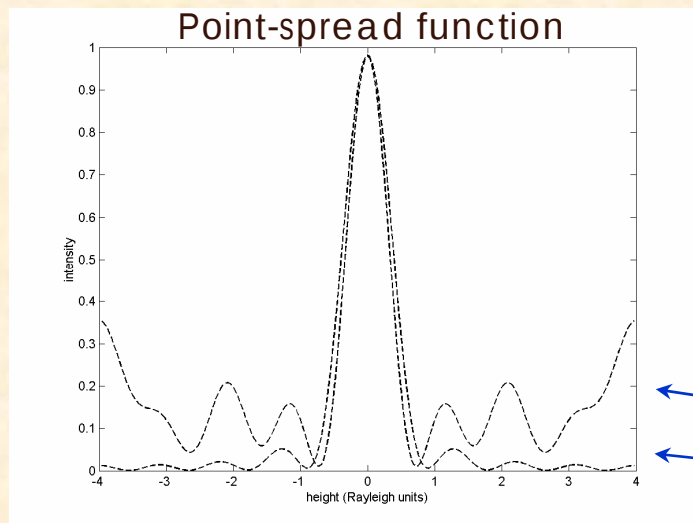
Wide field of potential applications, but:

Tradeoff ambiguity distance/z-resolution for given number K of tracks

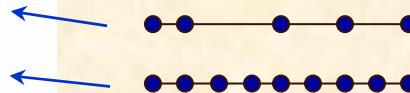
→ Limited z-resolution, many tracks required

Typically irregular track distribution:

→ **Anomalous sidelobes** in the PSF !
(quasi-grating lobes)



E-SAR, DLR

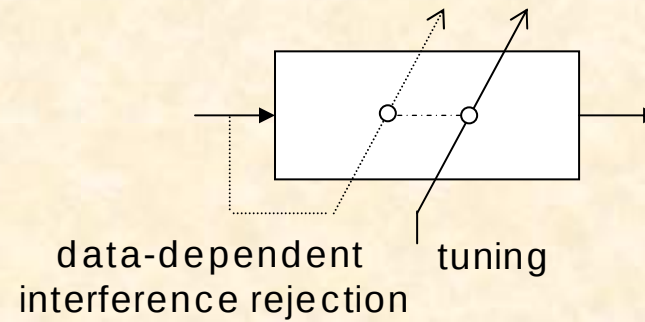
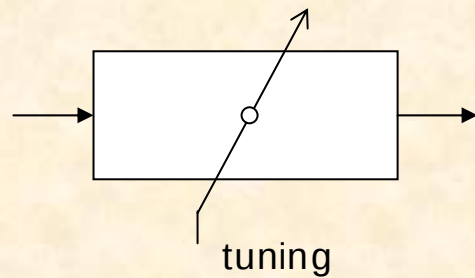


Stima spettrale non-parametrica *adattiva*

Lobi laterali → leakage → masking



Stima spettrale non-parametrica *adattiva*

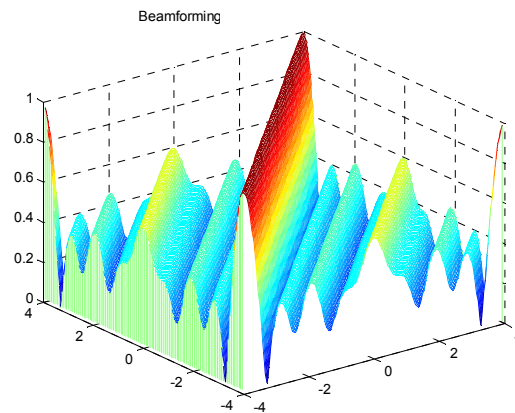
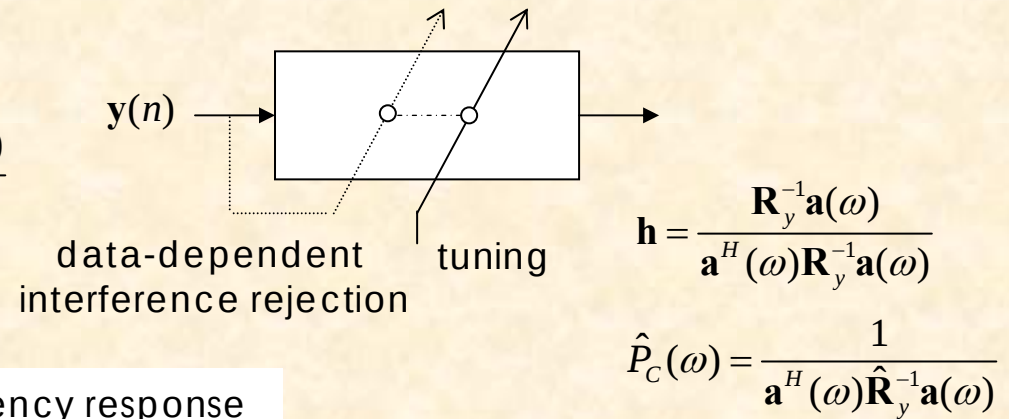
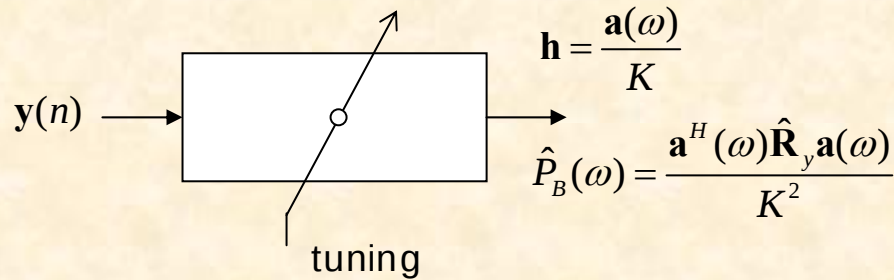


leakage suppression

Il filtro di Capon

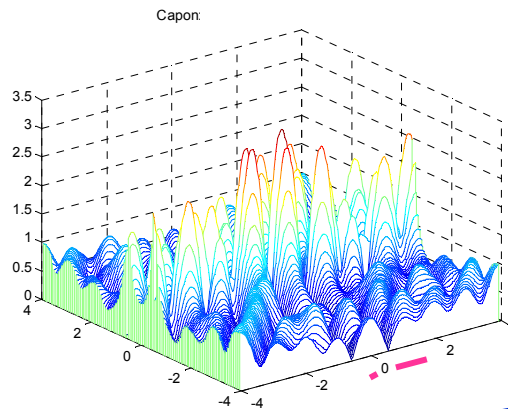
Stima MVU

Filtro sbiancante ad. + Period.



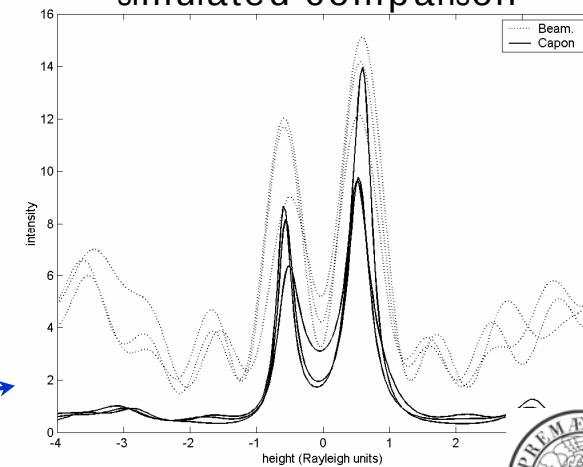
frequency response

$\bigcirc - \bigcirc - \bigcirc - \bigcirc - \bigcirc$
 $N=32$ looks SNR=9, 12 dB



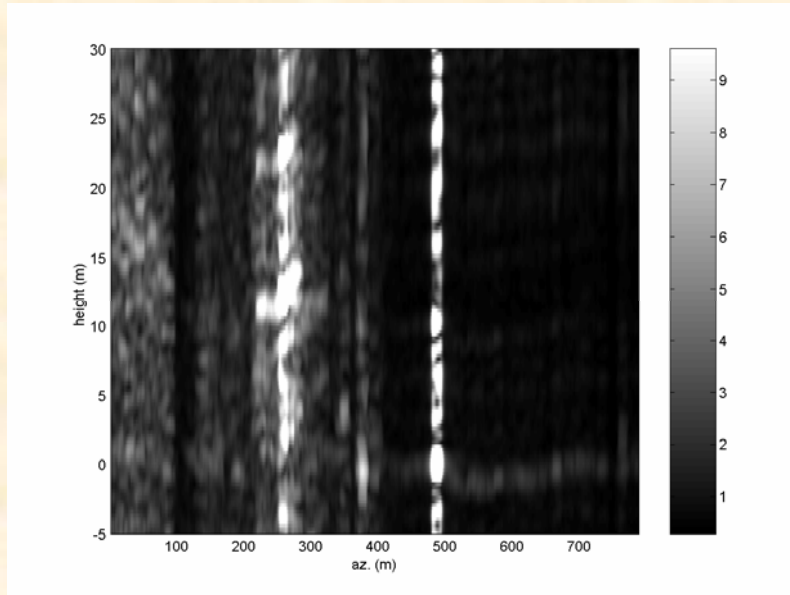
reduced leakage by adaptively setting nulls !

simulated comparison



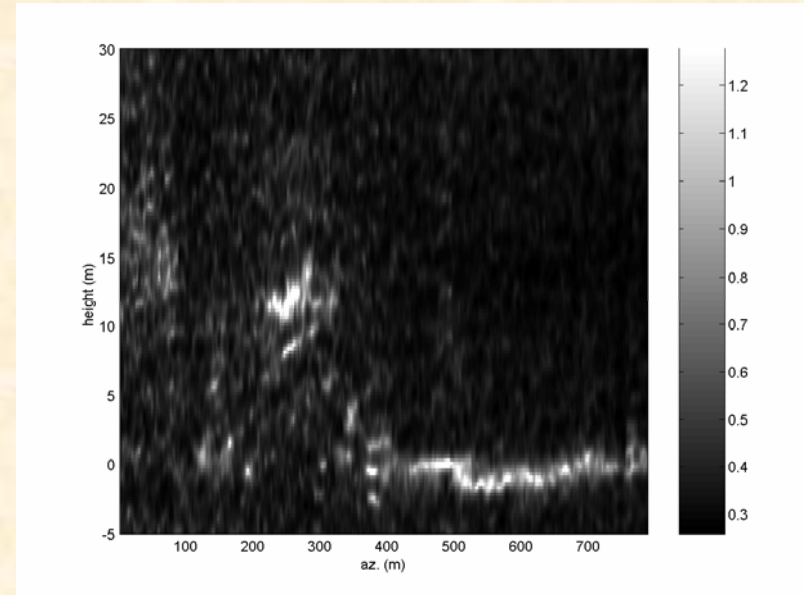
Radar d'immagine (Tomografici, Full 3D)

spatial Periodogram

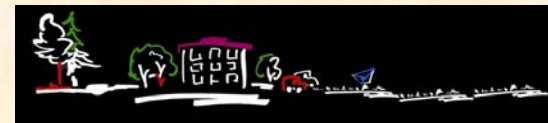


amplitude,
arbitrary units

Capon

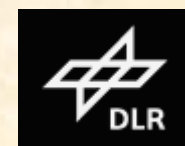
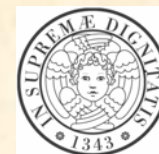


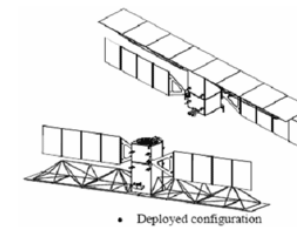
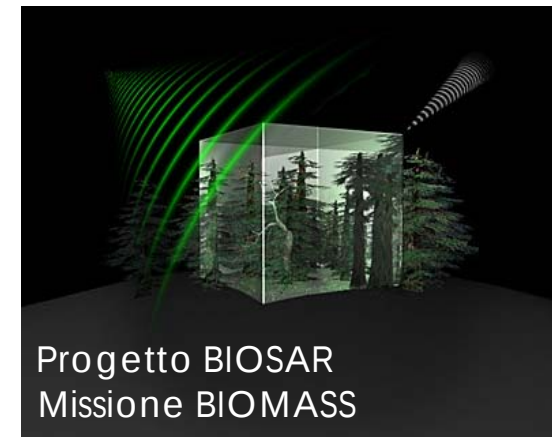
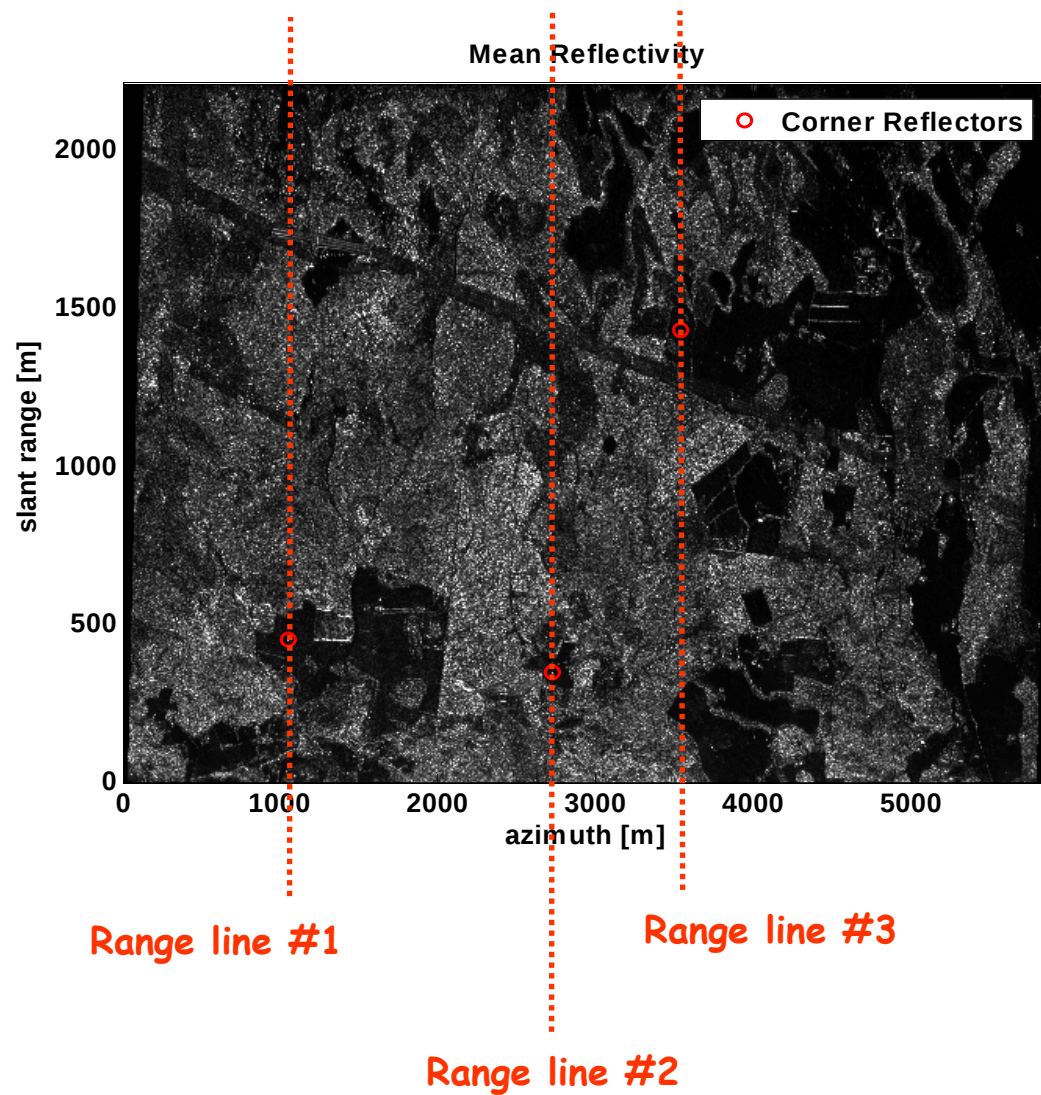
amplitude,
arbitrary units



K=14 tracks

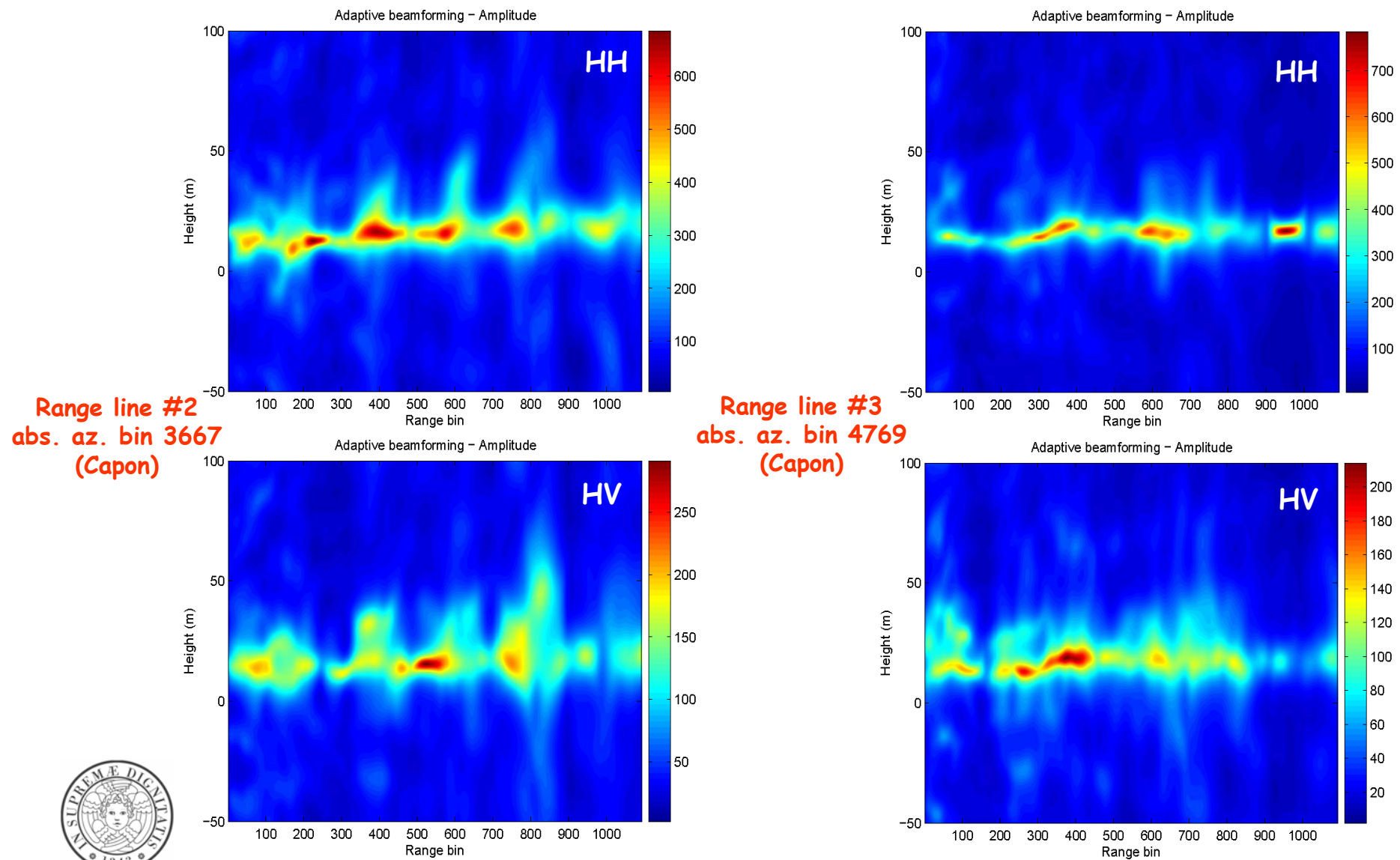
N=13 obs.







P-band data analysis





Differential Imaging

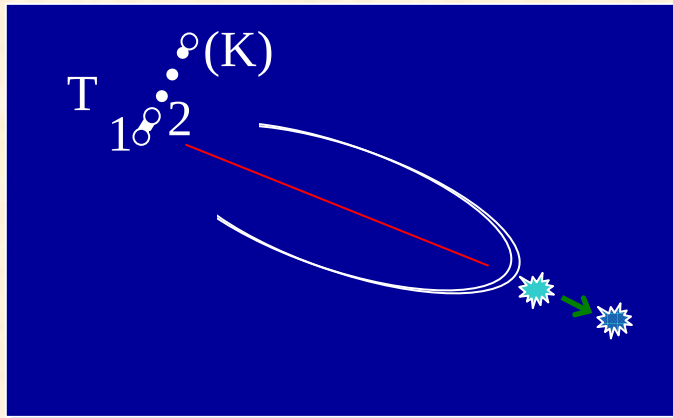
3D Tomo-SAR: multi-baseline

→ layover solution, h-backscatter profiling;
recently introduced

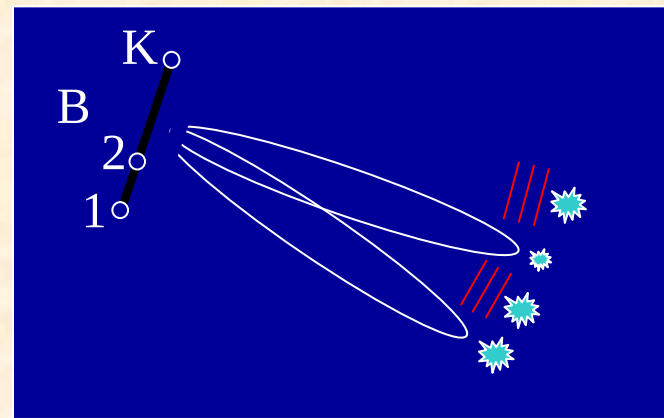
D-InSAR: multiple-pass

→ *displacements (single h-estimation);
operational technique*

“micro-Doppler” technique
(phase time changes: **temporal frequency**)



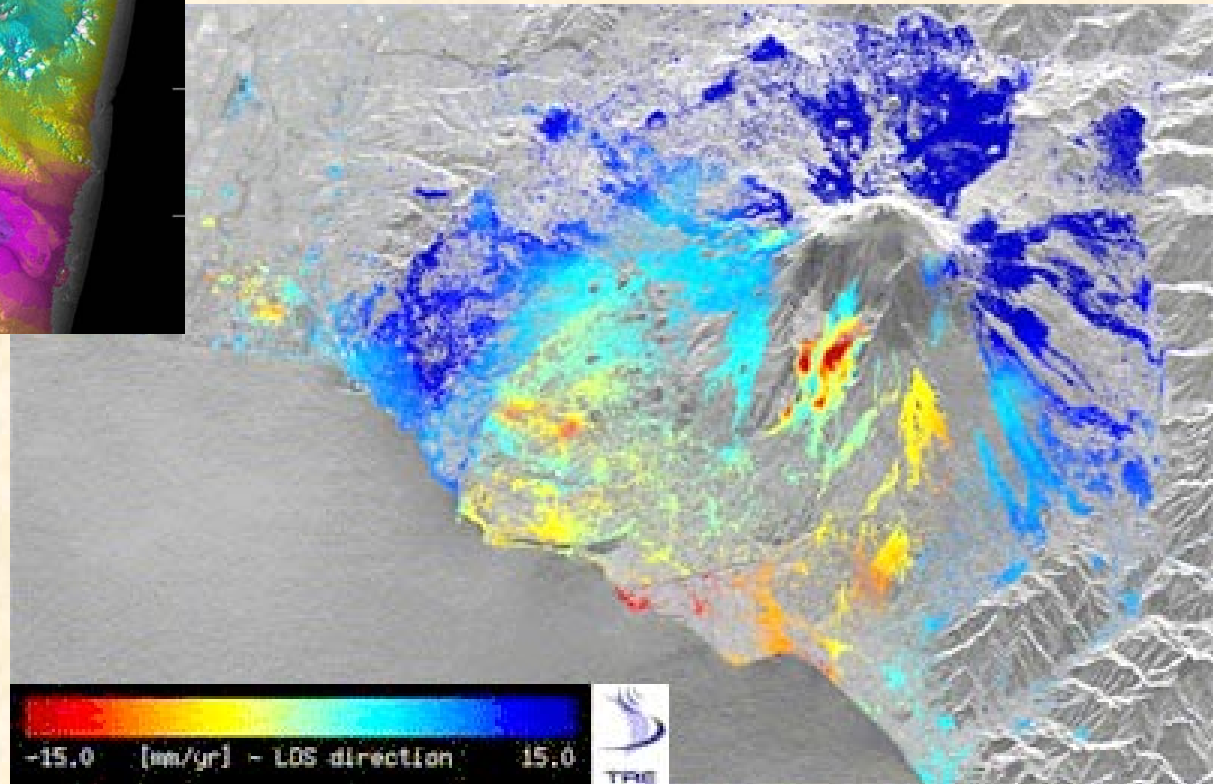
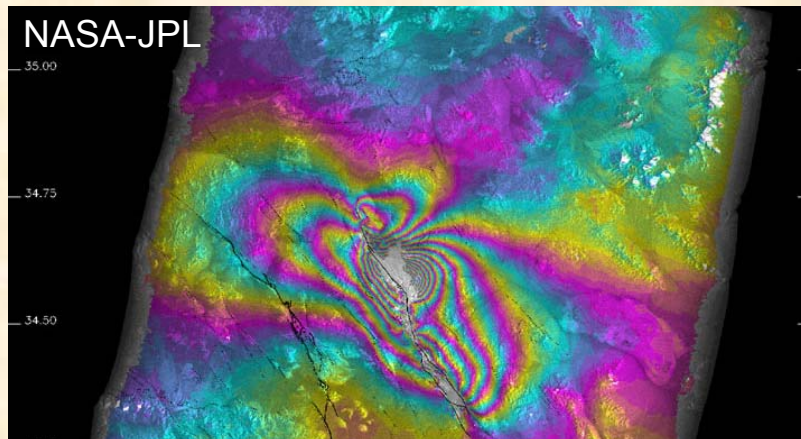
Beamforming technique
(**spatial frequencies**)



Radar d'immagine

(Tecniche Differenziali)

Applicazioni: analisi subsidenze, fronti di frana, faglie, terremoti, ...



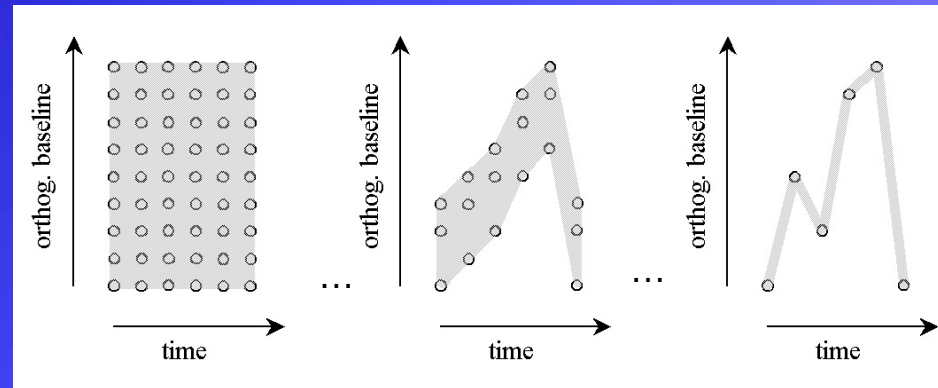
<http://eopi.esa.int/esa/esa?type=upload&ts=1071071849723&table=result&cmd=image&id=128>

“Differential Tomography”!



2D support in 2D Baseline-Time plane needed

Sparse 2D:
two-/multi-antenna
/sat. cluster
& multi-pass



“1D” Curvilinear:
plain multi-baseline
by multi-pass,
airborne/sat.

Processing Framework

after registration, compensation of phase artifacts, deramping:

$$\mathbf{A}(\omega_s, \omega_T) = \begin{bmatrix} 1 & \dots & e^{j[\omega_s B_{\perp}(1, N_P) + \omega_T t(N_P)]} \\ \vdots & \dots & \vdots \\ e^{j\omega_s} & \dots & \dots & \vdots(N_P) \end{bmatrix}$$

Bidimensional Baseline-Time
Spectral Estimation !
 $\mathbf{a}(\omega_s, \omega_T)$

→ z (3D) + Doppler resolution cell

Challenge: **Sparse Sampling** of 2D baseline-time support -
2D z-Doppler sidelobe problems to be attacked !

“4D Imaging”! (3D + T)

Simulated Bonn Data Set

ERS-1

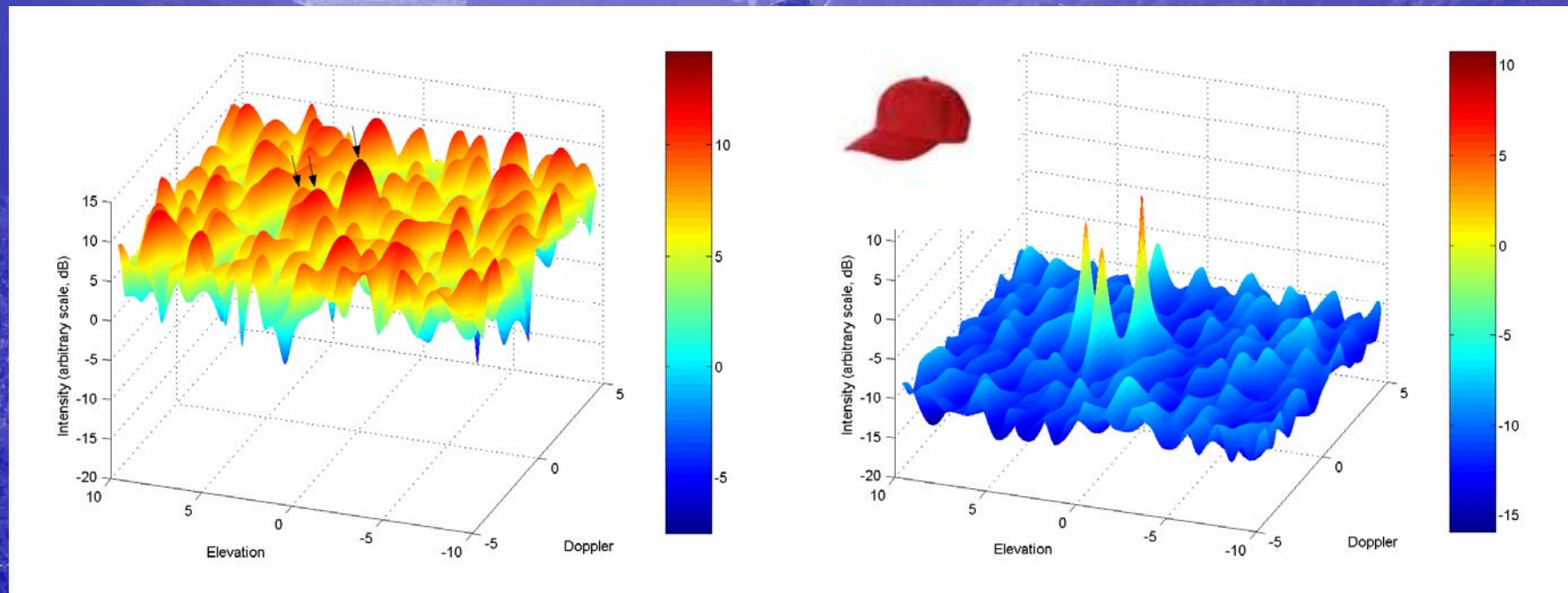
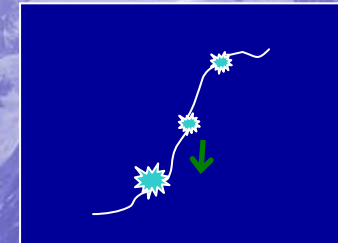
10 passes

time span: 27 days

overall baseline: 1418 m

SNR=15, 12, 9 dB

negligible baseline
and temporal decorrelation;
monotonic motion jitter.

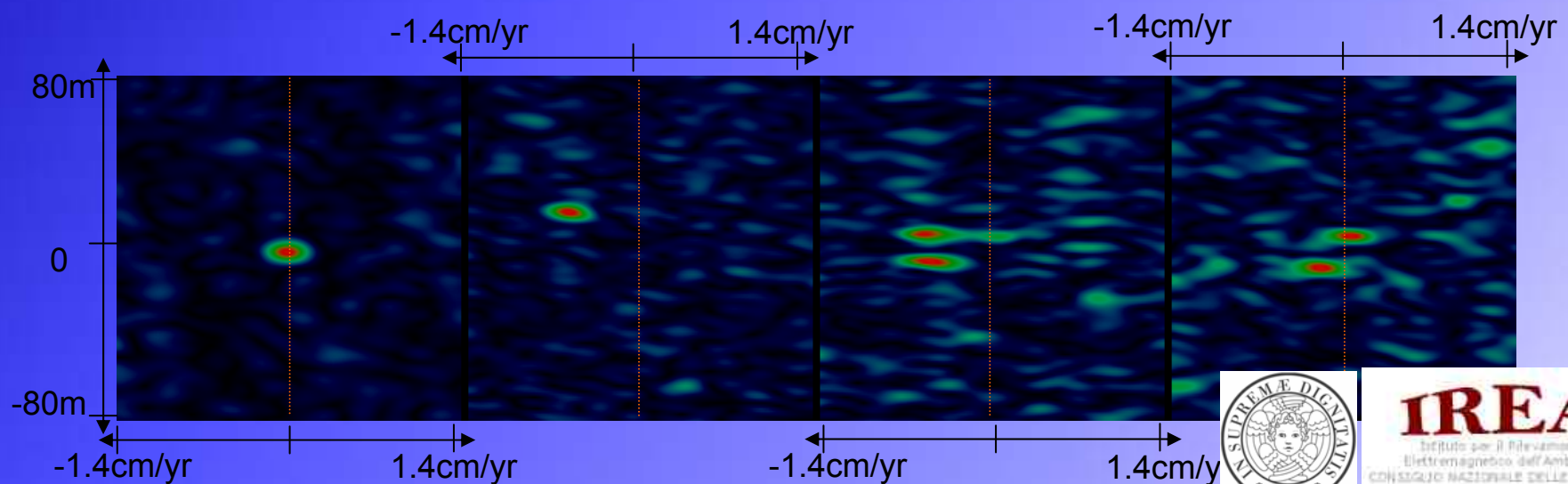
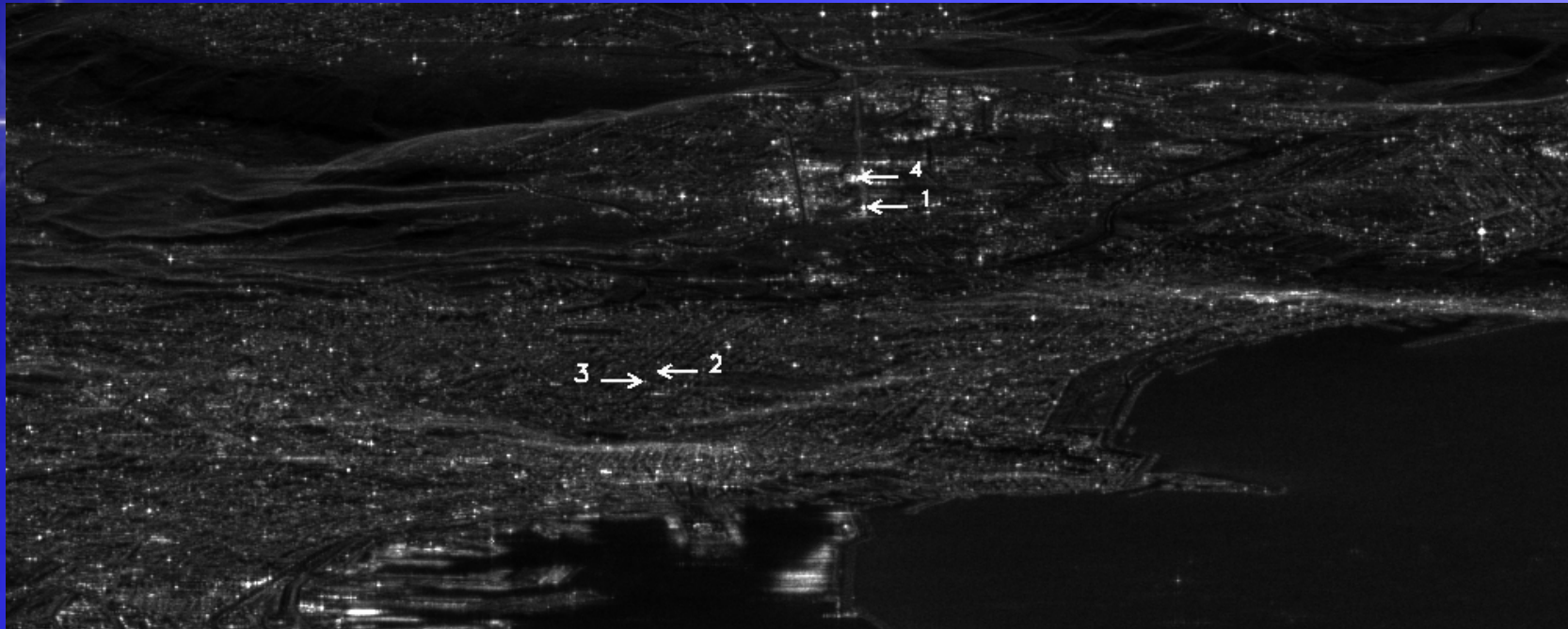


portion of Diff. + Tomo image

portion of Diff. + Tomo image



Results on real data of the ERS 1/2



Results on real data of the ERS 1/2



- Single scatterers -

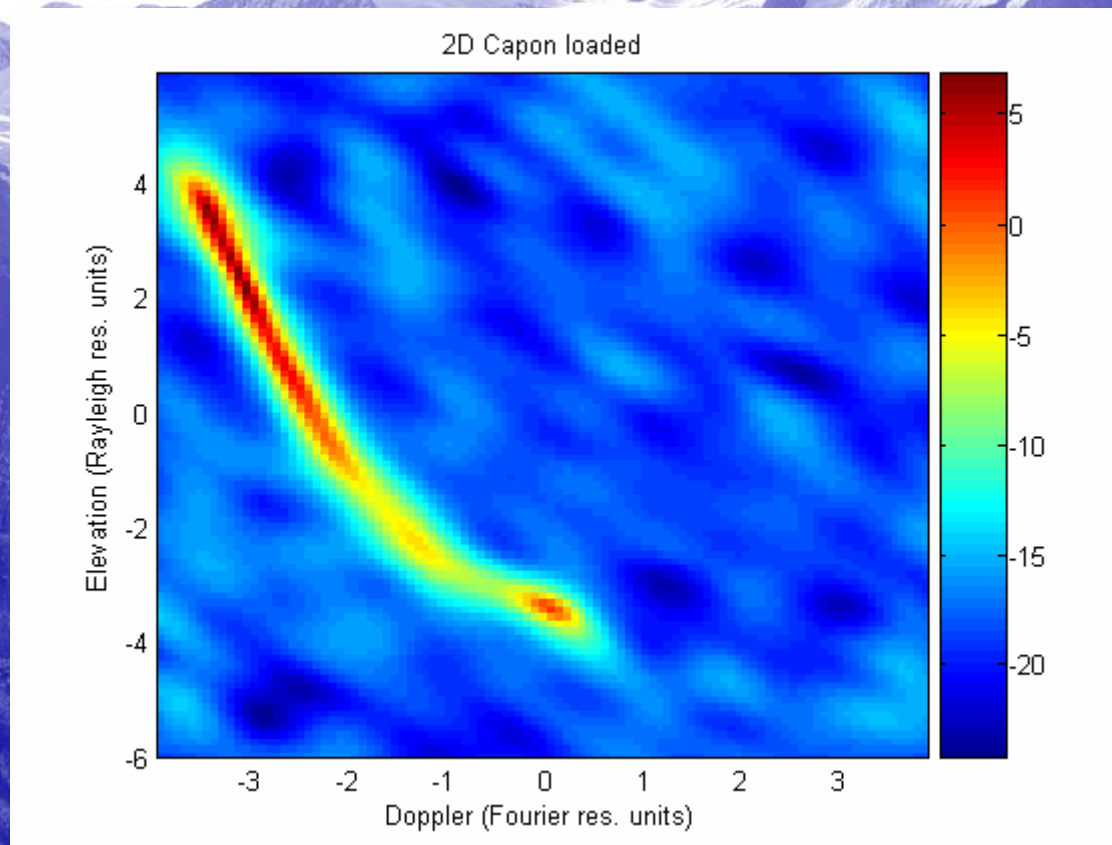


- Double scatterers -



“full 4D Imaging”! (continuous 3D + T)

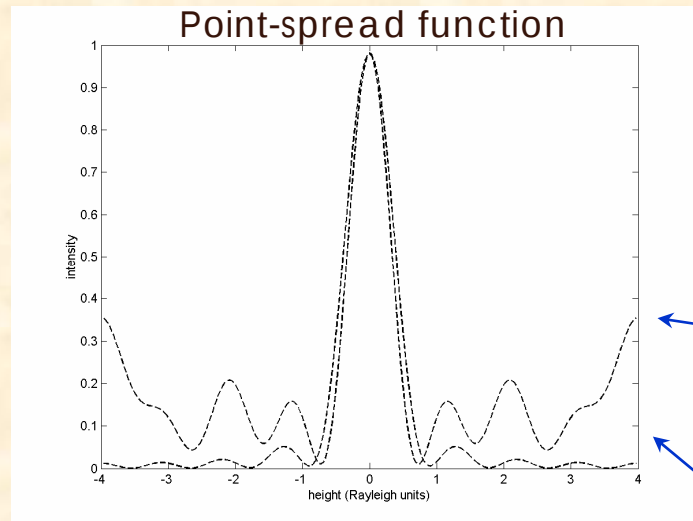
Simulated Results



“Eco-Doppler” radar imaging of a glacier flow !



Interpolation of non-uniformly sampled data

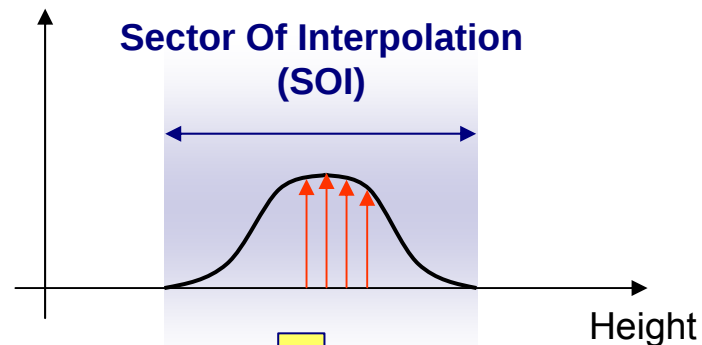


A-priori information:
spectral support
(Sector of Interpolation)

Problema di “predizione”

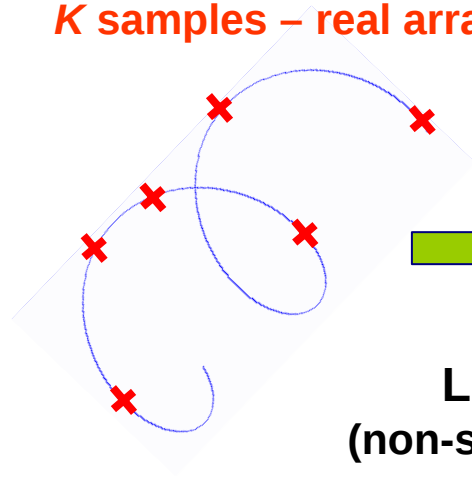


The new interpolated approach

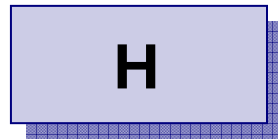


IDEA:
Reconstruct K_v uniform samples
of the spatial harmonic of interest
from the actual non-uniform samples

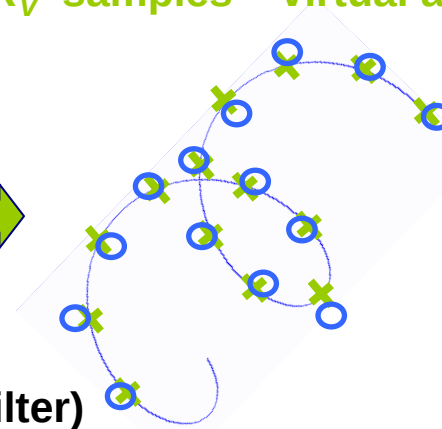
Non-uniform spatial sampling
 K samples – real array



Linear Interpolator
(non-stationary “matrix” filter)



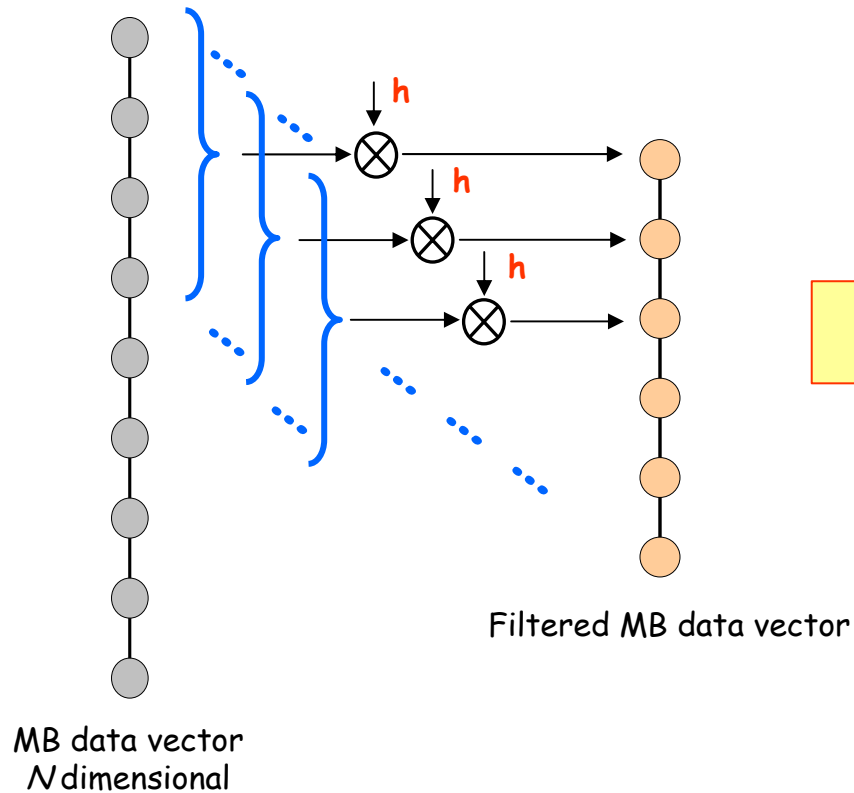
Uniform spatial sampling
 K_v samples – virtual array



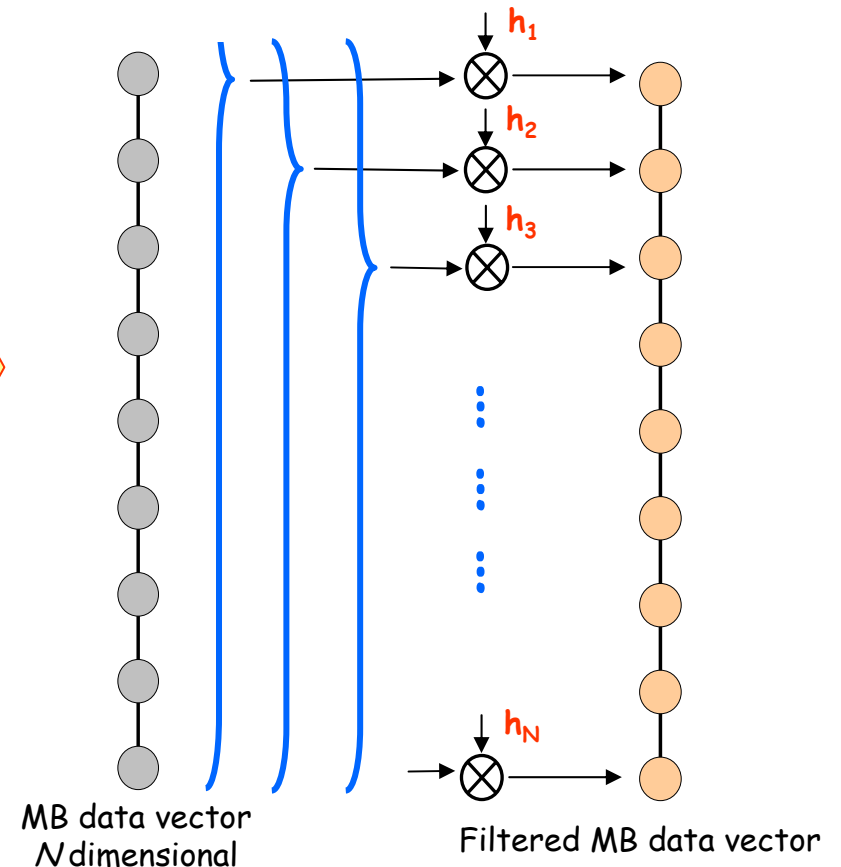
*Apply this process
for all the spatial
harmonics inside
the SOI!*

Matrix filter principle

Moving window filter - One coefficient vector h
Linear, spatially stationary



Matrix filter - Coefficient vectors $h_1, \dots, h_N$
Linear, spatially non-stationary



$\otimes$
inner product

• Arbitrary input/output MB spacing allowed

Array interpolation

Deterministic approach – IA

The output vector of the virtual array can be obtained by means of a $K_V \times K$ transformation matrix by solving the least squares (LS) problem:

$$\mathbf{H}_F = \arg \min_{\mathbf{H}} \|\mathbf{A}_V - \mathbf{H}\mathbf{A}\|_F^2 \quad \text{where}$$

$$\begin{aligned} \mathbf{A}_V &= [\mathbf{a}_V(\phi_1) \quad \mathbf{a}_V(\phi_2) \quad \cdots \quad \mathbf{a}_V(\phi_s)] \\ \mathbf{A} &= [\mathbf{a}(\phi_1) \quad \mathbf{a}(\phi_2) \quad \cdots \quad \mathbf{a}(\phi_s)] \end{aligned}$$

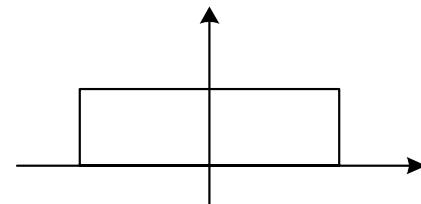
$[\phi_1 \phi_2 \dots \phi_s]$ are s phase samples inside the discretized SOI.

Statistical approach – MSE-IA

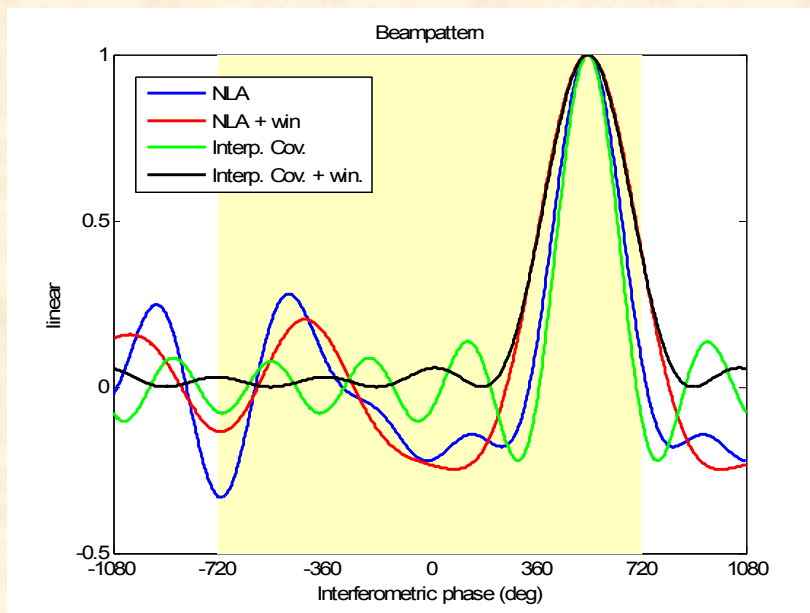
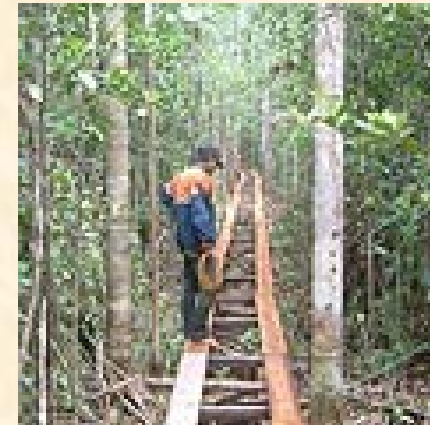
The signal vector χ_{ULA} can be estimated minimizing the mean square error (MSE):

$$\mathbf{H}_M = \arg \min_{\mathbf{H}} E \left\{ (\chi_{ULA} - \mathbf{H}\mathbf{y})^2 \right\}$$

This minimization (*Wiener-Hopf equations*) involves the knowledge of the spatial power spectral density, which can be assumed flat over the SOI in the InSAR application.

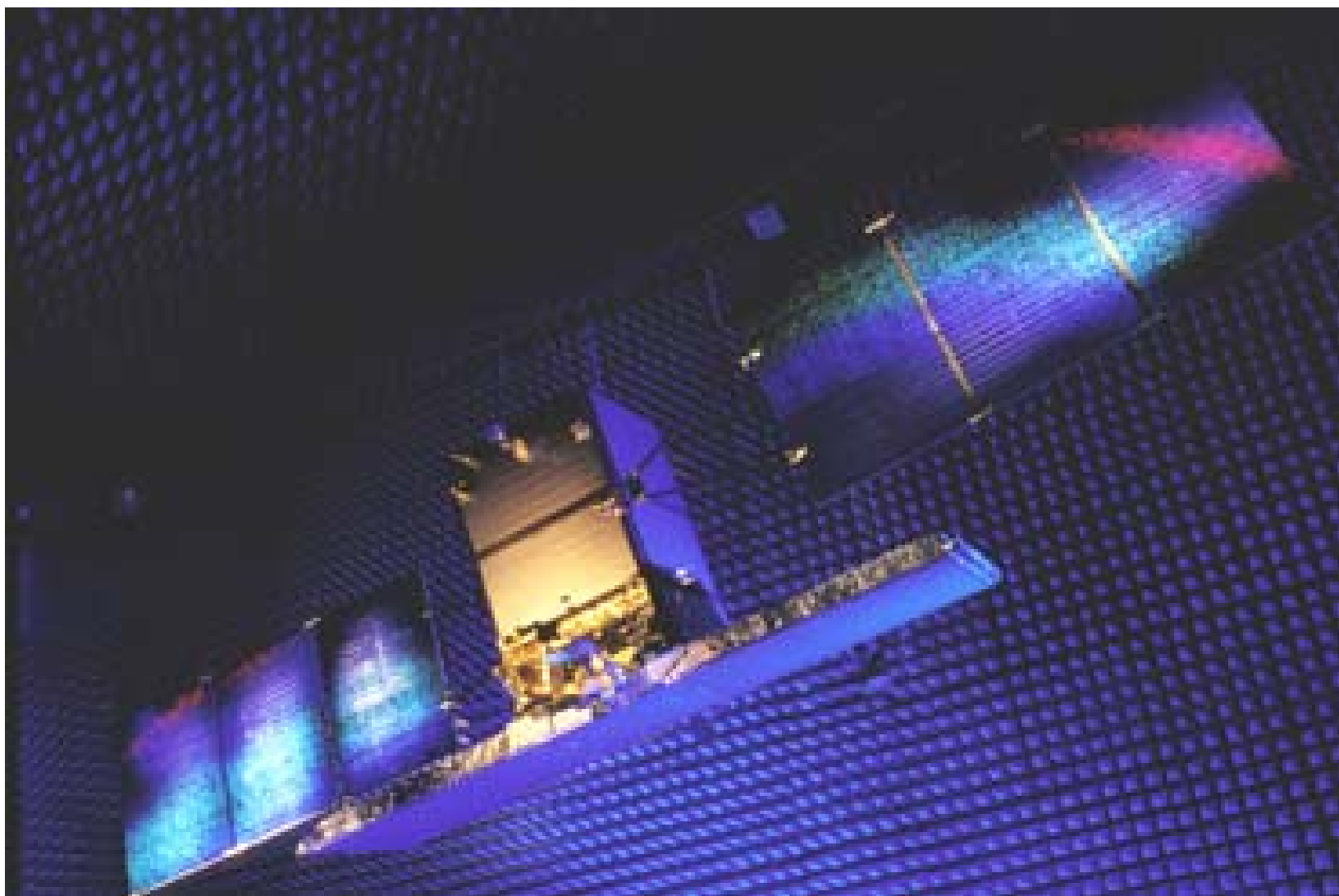


Costa Rica and Amazon Rainforest



Jet Propulsion Laboratory
California Institute of Technology





Il satellite italiano COSMO-SkyMed



Feasibility Study of Along-track SAR Interferometry with the COSMO-SkyMed Satellite System

Dept. of Ingegneria dell'Informazione, University of Pisa, via Caruso 14 - Pisa, Italy

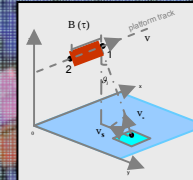


The Italian COSMO-SkyMed system

- 4-satellite constellation
- mapping of Mediterranean latitudes
- Hi-res X-band SAR
- Short revisit time

Along-Track Interferometry

- ocean surface current velocity
- ocean coherence time
- sea waveheight spectrum and MTI



Spaceborne ATI

- NASA SRTM (along-track boom component)
- TerraSAR-X split-antenna
- RADARSAT-II (split-antenna MTI)
- Satellite formations (distributed interferometers)

Performance study of *split-antenna* ATI modes for COSMO-SkyMed SAR:

Statistical data model

multilook data vector (up to 5 rx channels: Multi Baseline ATI)
 $y(n) = \sqrt{\sigma^2} \mathbf{A}(\varphi) \mathbf{x}(n) + \mathbf{v}(n)$

speckle autocorrelation function

$$[C_x]_{i,j} = \exp \left\{ - \left(\frac{(i-j)\tau}{(K-1)\tau_c} \right)^2 \right\}$$

data covariance matrix

$$\mathbf{R} = E\{y(n)y^H(n)\} = \sigma^2 \mathbf{A}(\varphi) \mathbf{C}_x \mathbf{A}^H(\varphi) + \sigma_v^2 \mathbf{I}$$

Cramér-Rao Lower Bound accuracy analysis

unknown and nuisance parameters $\chi = [\varphi \ \tau_c \ \sigma^2 \ \sigma_v^2]^T$

$\text{CRLB}(\chi_i) = [\mathbf{J}_{FIM}^{-1}]_{i,i}$ Fisher information

$$[\mathbf{J}_{FIM}]_{i,j} = N \text{tr} \left\{ \mathbf{R}^{-1} \frac{\partial \mathbf{R}}{\partial \chi_i} \mathbf{R}^{-1} \frac{\partial \mathbf{R}}{\partial \chi_j} \right\}$$

surface current velocity

$$v_s = \frac{\lambda v}{4\pi B \sin \vartheta_i} \varphi + v_B \quad \text{Bragg velocity}$$

accuracy of conventional 2-ch ATI

$$\text{CRLB}^{1/2}(v_s) = \frac{\lambda v}{4\pi B \sin \vartheta_i} \frac{1}{\sqrt{2N}} \frac{1}{\rho_s \rho_t} \sqrt{1 - (\rho_s \rho_t)^2}$$

System and physical models

SAR equation for extended targets, split antenna

Moore and Fung ocean scattering model

$$\sigma^0(\phi_{az}, \vartheta_i, u_w) = [A + B \cos \phi_{az} + C \cos 2\phi_{az}] e^{-(\vartheta_i - \bar{\vartheta})/\vartheta_0} \left(\frac{u_w}{u_w^*} \right)^\gamma$$

coherence time model (Frasier and Camps)

$$\tau_c = 3 \frac{\lambda}{u_w} \text{erf}^{-1/2} \left(2.7 \frac{r_g}{u_w^*} \right)$$

Platform velocity	7548 m/s
Slant range distance	732 Km
Antenna length	5.6 m
Carrier frequency	9.58 GHz
Incidence angle	25°-50°
TX peak power	5 KW
Bandwidth	99.3 MHz

A few split mode configurations:	# rx panels per ch.	# ch.	B [m]
STTX+SPAN2a	1	2	2.24
STTX+SPAN2b	2	2	1.68
STTX+SPAN2c	3	2	1.12
STTX+SPAN3d	3	3	1.12

Rankings

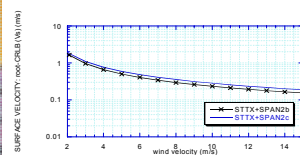
1° STTX+SPAN4/SPAN5
2° STTX +SPAN2b
3° STTX +SPAN2a/SPAN3a
4° STTX +SPAN2c/3d

Conv. and MB coherence time estimation and quantization effects also investigated

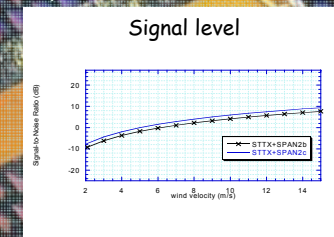
Investigated trade-offs

- baseline length
- sub-aperture gain
- channel number
- speckle decorrelation azimuth ambiguity

Velocity accuracy

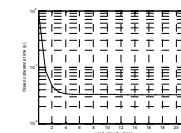


A few results:



cross-wind, mid-range, $r_g \approx 2.7$ m

Coherence time



up- and down-wind

Resolution

	Accuracy=10 cm/s, mid-range	Accuracy=25 cm/s, mid-range	Accuracy=25 cm/s, far-range
STTX+SPAN2b	N=189035 ATI cell=1271 m	N=30245 ATI cell=508 m	N=50592 ATI cell=657 m
STTX+SPAN2c	N=235683 ATI cell=1419 m	N=37709 ATI cell=568 m	N=58602 ATI cell=708 m

cross-wind



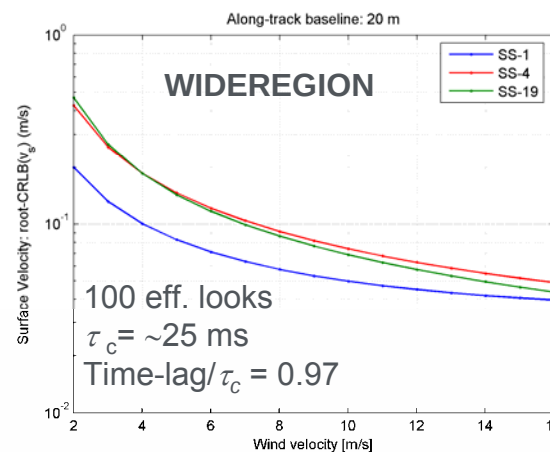
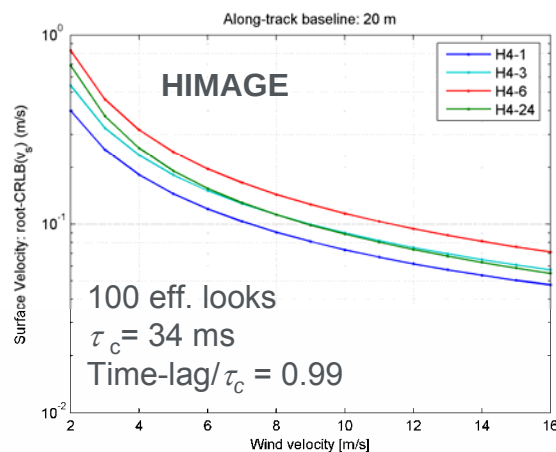
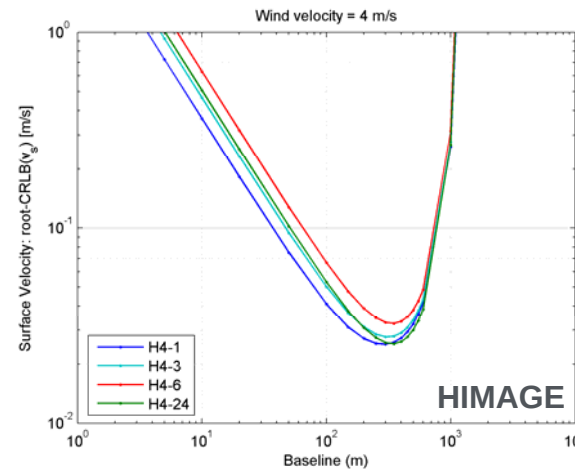
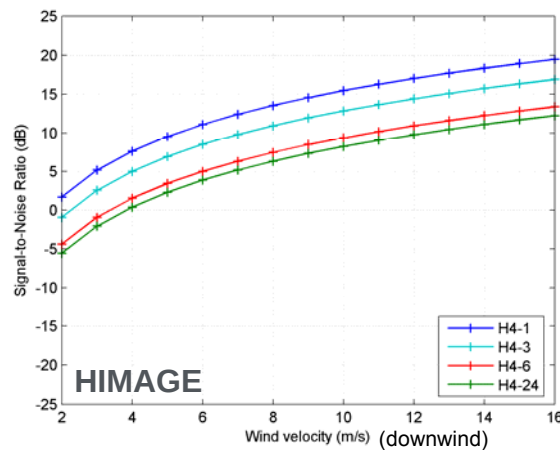
Lancio del satellite italiano COSMO-SkyMed (COSMO-1, Giugno 2007)

COSMO-SkyMed/BISSAT cluster

ATI Sea Currents and Waves



- Cramér-Rao bound analysis on the expected precision on surface current velocity estimation
- SAR equation for extended target
- Physical models for ocean scattering and coherence time (τ_c)



- A rescaling of the number of looks is evaluated to obtain the desired precision (≤ 0.1 m/s at worst)
- @ $B_{ATI} = 20$ m, ~ 0.1 m/s absolute surface velocity accuracy is obtained with $B_{XTI} < 500$ m (availability of precise radar altimeter oceanographic data and DEM of a terrain zone in the image)

Possible extension: BISSAT with COSMO-SkyMed in split-antenna mode

- Dual baseline system
- Small baseline: enlargement of the velocity ambiguous range
- Long baseline: augmented interferometric sensitivity, eased absolute calibration

Example: 200 m total baseline, 3.36 m small baseline (STTX-SPAN2b): higher multilooking degree (50 instead of 15 in the single baseline case) to ensure the desired precision of 0.1 m/s

System model perturbations

Residual multibaseline array calibration or atmospheric compensation errors affect the data in all the practical applications

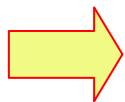
Phase error vector introduced by the residual miscalibration at each array phase center:

$$\mathbf{e} = [e^{j4\pi d_1} \quad e^{j4\pi d_2} \quad \dots \quad e^{j4\pi d_K}]^T$$

$d_1 \dots d_K$

Gaussian, zero mean, uncorrelated random variables, with variance σ^2

- **Positioning errors** of the array elements (normalized to wavelength), due to uncertainties in the platform positioning in the direction parallel to the line of sight (errors in the perpendicular direction can be neglected)
- **Residual one-way path delay** variations of wave propagation through the atmosphere (normalized to wavelength)



Total array response: $\mathbf{a}(h) = \mathbf{a}_P(h) \odot \mathbf{e}$

Element-by-element product

Hybrid Cramér-Rao Bound

Signal model:

$K = \#$ of array elements
 $N_s = \#$ of sources

$$\mathbf{y}(n) \approx \sum_{i=1}^{N_s} \sqrt{\tau_i} \mathbf{a}_p(\varphi_i) \odot \mathbf{e} \odot \mathbf{x}_i(n) + \mathbf{v}(n) \quad n=1,2,\dots,N$$

Speckle vector (multiplicative noise)

Steering vector *Calibration error vector*

Cramér-Rao bound

$$\xi = [\varphi_1 \quad \dots \quad \varphi_{N_s} \quad \tau_1 \quad \dots \quad \tau_{N_s} \quad b_1 \quad \dots \quad b_{N_s} \quad \sigma_v^2]^T \quad \text{Parameter vector}$$

$$[\mathbf{J}_{FIM}]_{m,n} = -\mathbb{E}_{\mathbf{y}} \left\{ \frac{\partial^2 \ln \left(\iint \dots \int f(\mathbf{y}, \mathbf{e}; \xi) d\mathbf{d}_1 d\mathbf{d}_2 \dots d\mathbf{d}_K \right)}{\partial \xi_m \partial \xi_n} \right\} \Rightarrow \text{Unfeasible !}$$

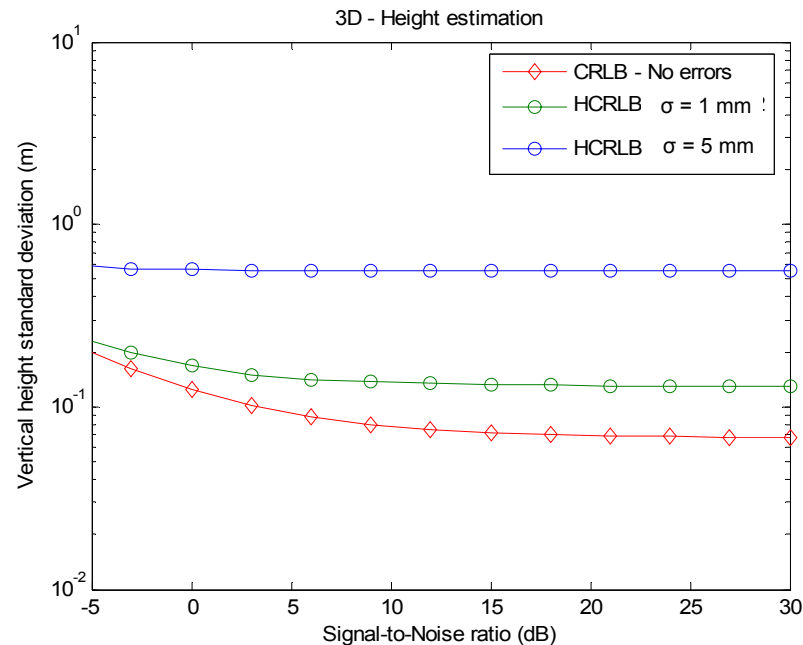
Hybrid Cramér-Rao bound

Good trade-off between easiness of calculation and tightness

$$\xi = [\varphi_1 \quad \dots \quad \varphi_{N_s} \quad \tau_1 \quad \dots \quad \tau_{N_s} \quad b_1 \quad \dots \quad b_{N_s} \quad \sigma_v^2 \quad d_1 \quad \dots \quad d_K]^T \quad \text{Nuisance parameters}$$

$$[\mathbf{J}_{H-FIM}]_{m,n} = -\mathbb{E}_{\mathbf{y}, \mathbf{d}} \left\{ \frac{\partial^2 \ln f(\mathbf{y}, \mathbf{e}; \xi)}{\partial \xi_m \partial \xi_n} \right\} = -\mathbb{E}_{\mathbf{d}} \left\{ \mathbb{E}_{\mathbf{y}|\mathbf{d}} \left\{ \frac{\partial^2 \ln f(\mathbf{y}, \mathbf{e}; \xi)}{\partial \xi_m \partial \xi_n} \right\} \right\} \Rightarrow HCRB(\xi_k) = [\mathbf{J}_{H-FIM}^{-1}]_{k,k}$$

System model perturbations: *Hybrid CRLB*



3D: height

Double scatterer distance in height: 2 r.u.

Notes:

- σ is standard deviation of residual atmospheric one-way path delay variations
- Variations uncorrelated from track to track
- Compensated I.o.s. velocity



Robust interpolator derivation

Again: $\mathbf{y}_V(n) = \mathbf{H}_I \mathbf{y}(n)$

Robust interpolator, $K_V \times K$ matrix

In the perfectly calibrated case: $\mathbf{H}_I \mathbf{A}_P = \mathbf{A}_{P,V}$ *Overdetermined equation system:
LS solution*

NLA array steering matrix:
*The columns are the NLA st.
vectors calculated for a number of
spatial frequency in the SOI*

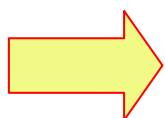
Virtual array steering matrix:
*The columns are the virtual st.
vectors calculated for the same
spatial frequencies in the SOI*

In the presence of calibration errors:

$$\mathbf{H}_I (\mathbf{A}_P + \Delta \mathbf{A}_P) = \mathbf{A}_{P,V} + \Delta \mathbf{A}_{P,V}$$

Calibration error

Interpolation error

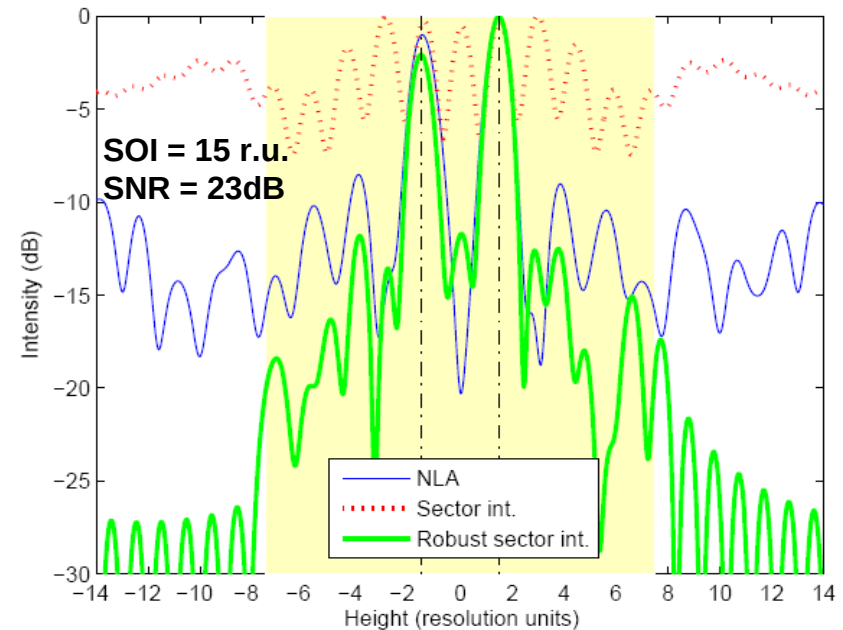
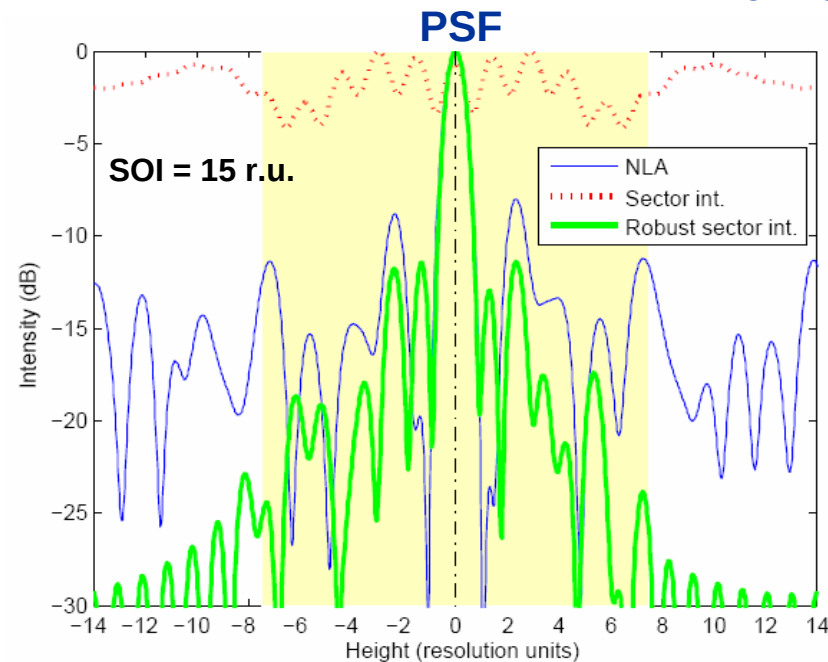


$$\mathbf{H}_I = \arg \min_{\mathbf{H}} E \left\{ \|\Delta \mathbf{A}_{P,V}\|_F^2 \right\}$$

*LS solution to a
perturbed equation system*

Results of the robust processing

$K = 24$ highly non uniform, $K_v = 29$
 $\sigma = 0.01$ (λ units)



- The basic sector interpolator degrades in the reconstruction of the tomographic profile
- The higher effectiveness of the robust method is apparent